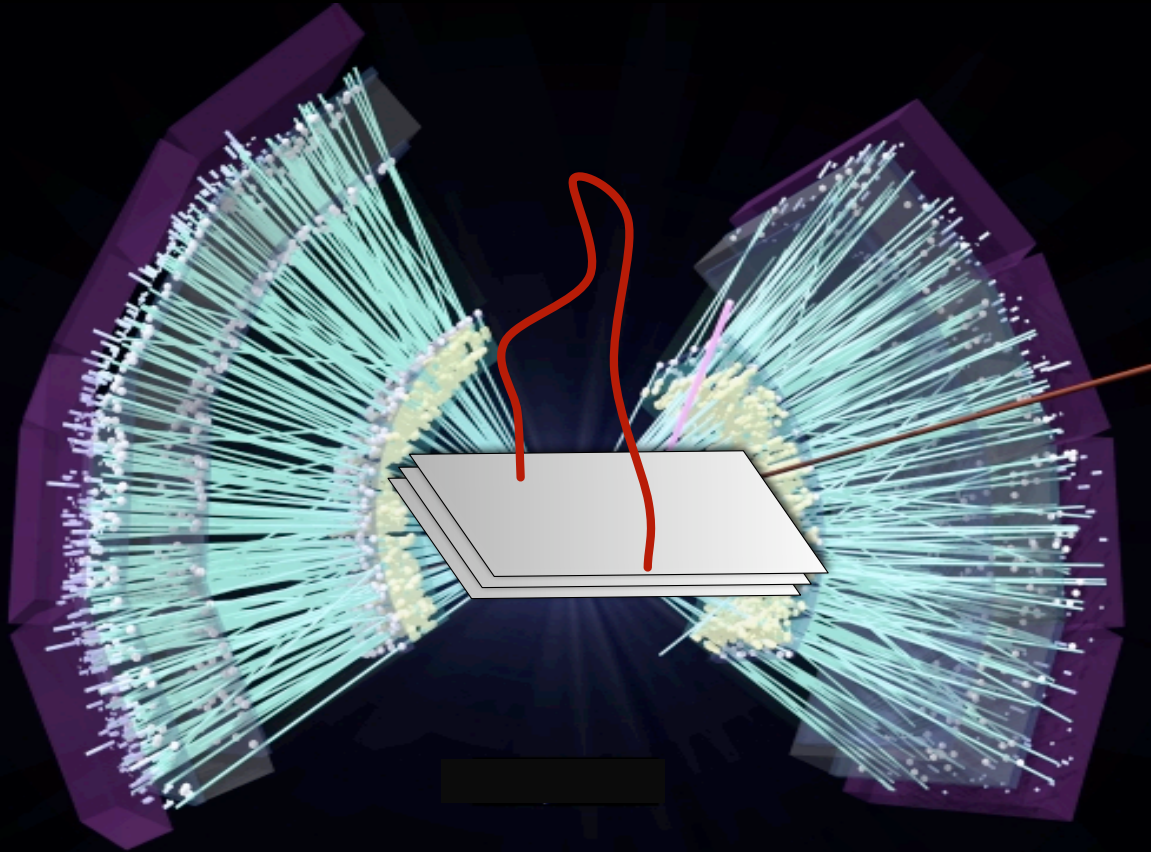


RHIC physics and String Theory: The Fundamental Story



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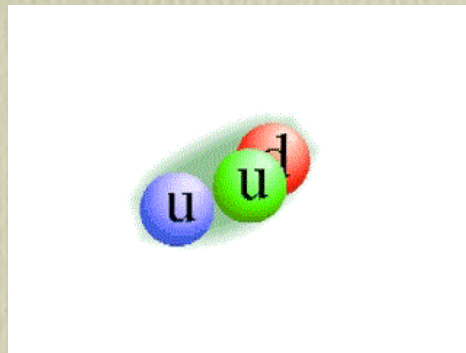
Plan

- i) Motivation.
- ii) Phase transitions.
- iii) Viscosity.
- iv) Photon production.
- v) Finite density.
- vi) Future prospects.

i) Motivation.

The QCD challenge

- QCD remains a challenge after 34 years!



The QCD challenge

- QCD remains a challenge after 34 years!
- A string reformulation might help.
- Lots of gauge/gravity examples.
- Unfortunately, QCD dual is not accessible via supergravity.

Therefore:

- Certain quantitative observables (eg. $T=0$ spectrum) will require going beyond supergravity.
- However, certain predictions may be universal enough to apply in certain regimes.
- Good example: $\eta/s = 1/4\pi$ Gubser, Klebanov & Peet '96
Policastro, Son & Starinets '01
Same for all non-Abelian plasmas with gravity dual!
- How about QCD just above deconfinement?



Results from RHIC indicate $\eta/s \sim 1/4\pi$.

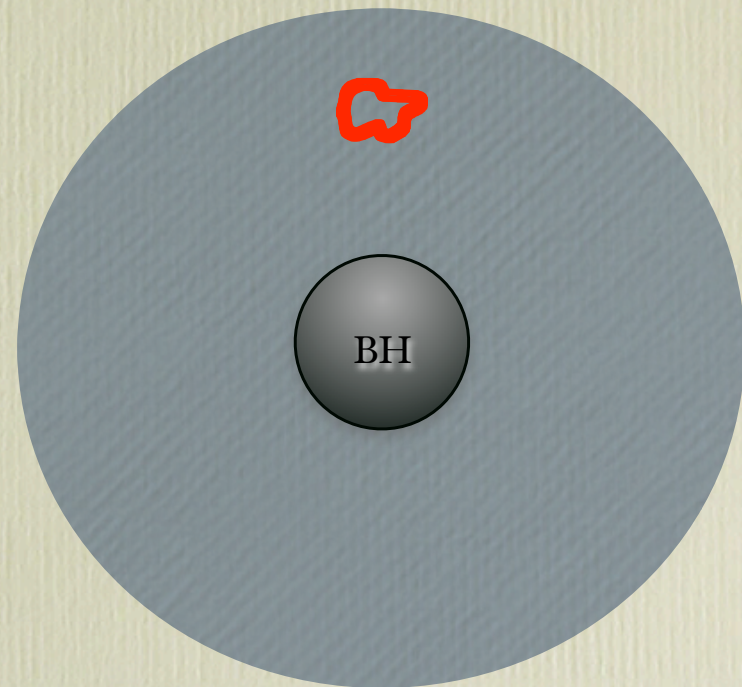
Animation by Jeffery Mitchell (Brookhaven National Laboratory). Simulation by the UrQMD Collaboration

Observations:

- Did not know η/s was going to be universal! Kovtun, Son & Starinets '03

- Based on universal property:

Gravity dual of a deconfined plasma contains a black hole



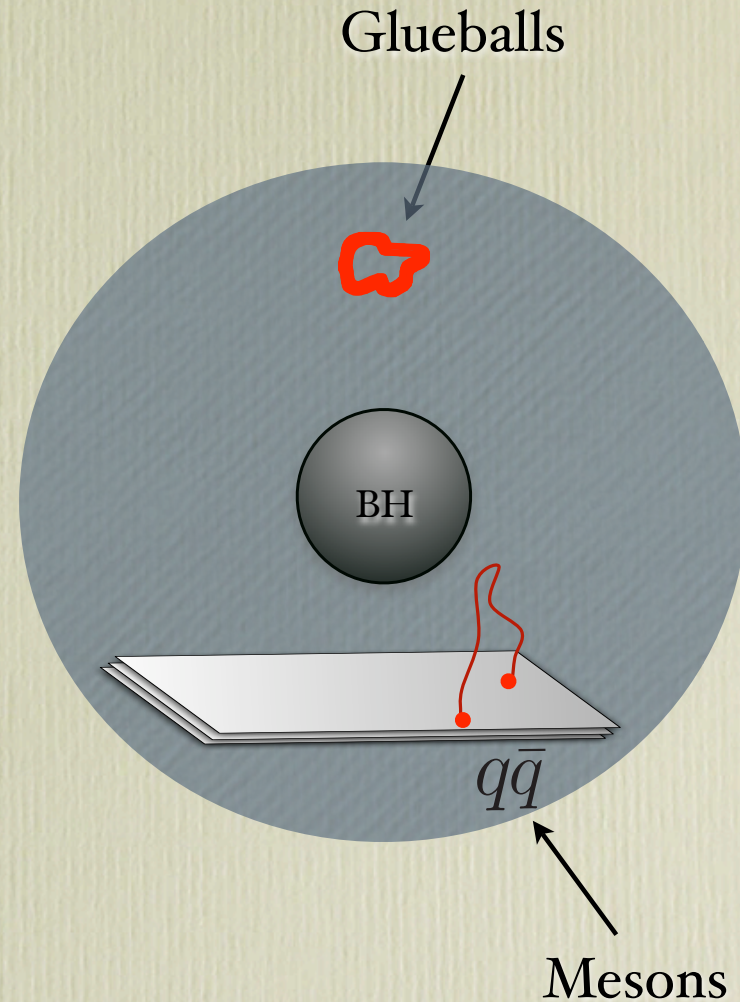
Observations:

- Combine with another one:

$N_f \ll N_c$ quark flavours correspond
to N_f probe branes

Karch & Katz '02

For concreteness will concentrate
on D7 probes in D3 background.



Disclaimer:

Not QCD, so interpret with caution.

ii) Fundamental phase transitions.

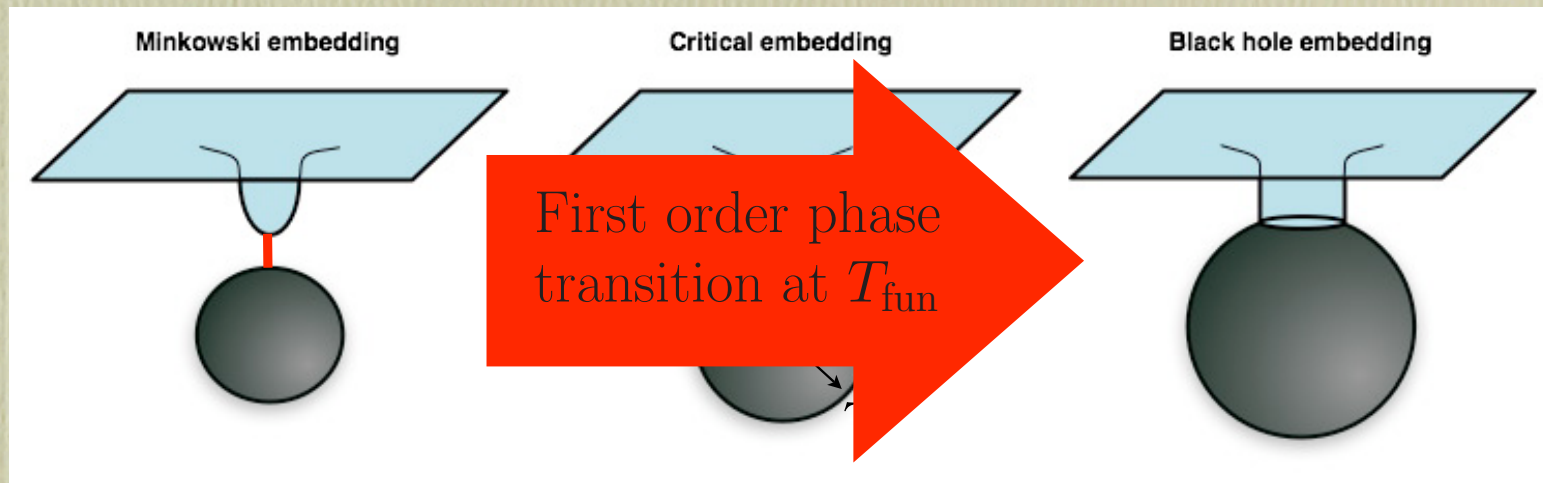
D.M., Myers & Thomson '06

Previous related work:

Babington, Erdmenger, Guralnik & Kirsch '03

Kruczenski, D.M., Myers & Winters '03

Kirsch '04



- Discrete set of mesons with mass gap:

$$M_{\text{mes}} \sim \frac{M_q}{\sqrt{\lambda}} \sim T_{\text{fun}}$$

- Free massive quarks.
- Absolutely stable -- survive deconfinement!
- In good agreement with lattice QCD, eg. for J/Ψ :

Lattice: $T_{\text{fun}} \simeq (317 - 403) \text{ MeV}$

Gravity: $T_{\text{fun}} \simeq (371 - 712) \text{ MeV}$

- No quasi-particle excitations!

iii) Viscosity of fundamental matter.

D.M., Myers & Thomson '06

Universal viscosity bound

- Conjectured universal bound for relativistic plasmas:

$$\frac{\eta}{s} \geq \frac{1}{4\pi}$$

Kovtun, Son & Starinets '03

- Saturated at $N_c \rightarrow \infty$, $\lambda \rightarrow \infty$ by all holographic theories with adjoint matter.

- Results from RHIC are close: $\frac{\eta}{s} \sim \frac{1}{4\pi}$

Teaney '03

Shuryak '03

Romatschke & Romatschke '07

Heinz '07

- What about when quarks are included?
QCD and leading N_f/N_c correction.

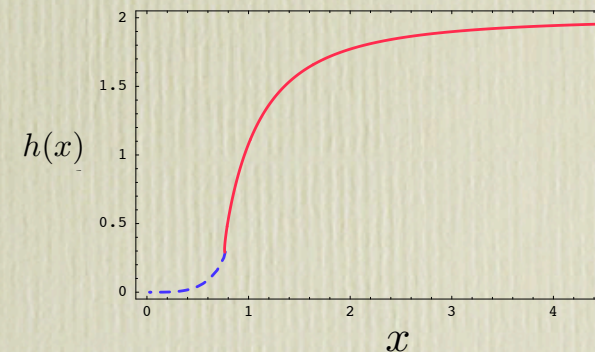
Holographic calculation

- Kubo: $\eta = \lim_{\omega \rightarrow 0} \frac{1}{2\omega} \int dt d^3x e^{i\omega t} \langle [T_{ij}(x), T_{ij}(0)] \rangle$

Couples to g_{ij} \longrightarrow $\eta \sim \frac{\delta^2 S_{D7}}{\delta g_{ij}^2}$

- η and S are corrected at $\mathcal{O}\left(\frac{N_f}{N_c}\right)$ but the ratio is not!

$$\eta = \frac{\pi}{8} N_c^2 T^3 \left[1 + \frac{\lambda}{8\pi^2} \frac{N_f}{N_c} h\left(\frac{\lambda T}{M_q}\right) + \dots \right]$$



Universal enhancement
with simple origin:

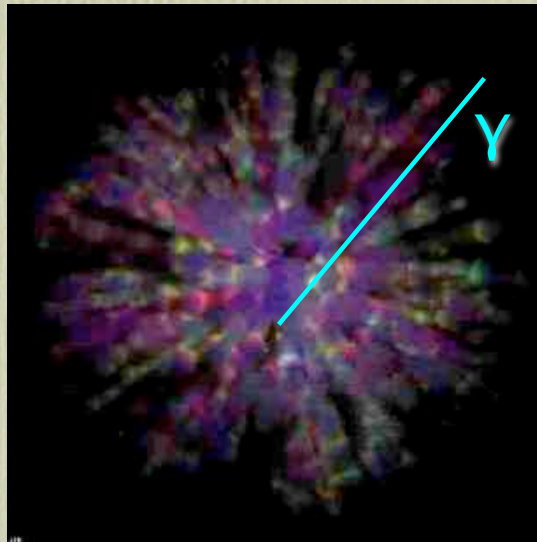
$$G_N T_{D7} \sim \frac{\lambda N_f}{N_c}$$

iv) Holographic photon production.

D.M., Patiño-Jaidar (to appear)

Why photons?

- QGP is optically thin \rightarrow Photons carry valuable information.



- Holographic results for massless matter:

Caron-Huot, Kovtun, Moore, Starinets & Yaffe '06

Parnachev & Sahakian '06

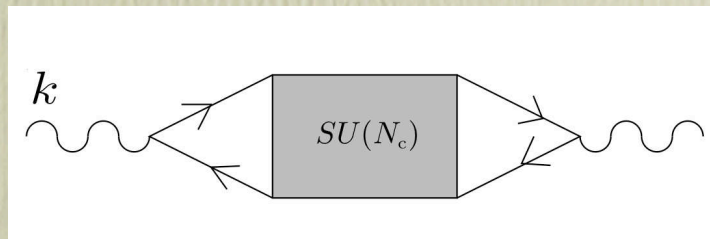
- To leading order in the electromagnetic coupling constant:

$$\frac{d\Gamma}{d^d\mathbf{k}} = \frac{e^2}{(2\pi)^d 2|\mathbf{k}|} \frac{1}{e^{k^0/T} - 1} \eta^{\mu\nu} \chi_{\mu\nu}(k)$$

$k = (k^0, \mathbf{k})$, with $k^0 = |\mathbf{k}|$, is the photon null momentum

$\chi_{\mu\nu}(k) = -2 \text{Im} G_{\mu\nu}^{\text{R}}(k)$ is the spectral density

$$G_{\mu\nu}^{\text{R}}(k) = -i \int d^{d+1}x e^{-ik \cdot x} \Theta(x^0) \langle [J_{\mu}^{\text{EM}}(x), J_{\nu}^{\text{EM}}(0)] \rangle$$



Holographic calculation

Gauge theory

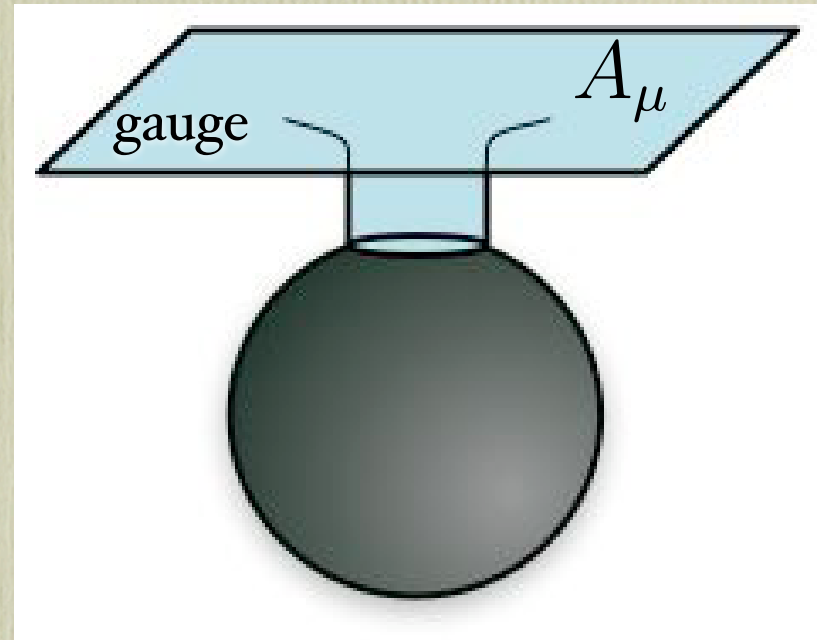
$$U(N_f) \simeq SU(N_f) \times U(1)_B$$

$$\text{Conserved } J_\mu^B = J_\mu^{\text{EM}}$$

AdS/CFT prescription:

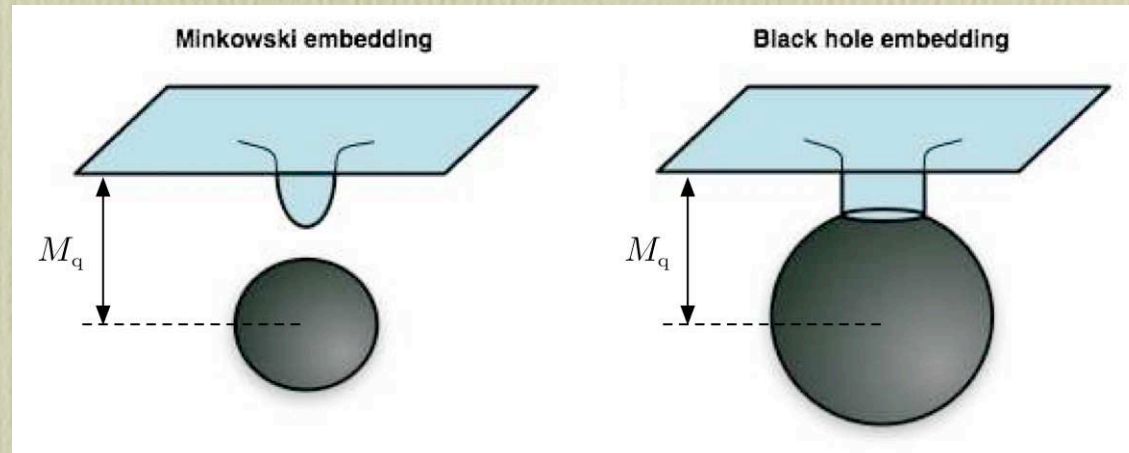
$$G_{\mu\nu}^{\text{R}} \sim \frac{\delta^2 S_{\text{D7}}}{\delta A_\mu \delta A_\nu}$$

String theory



Comments:

- Concentrate on BH embeddings:



$$\chi = \sum \text{delta functions}$$

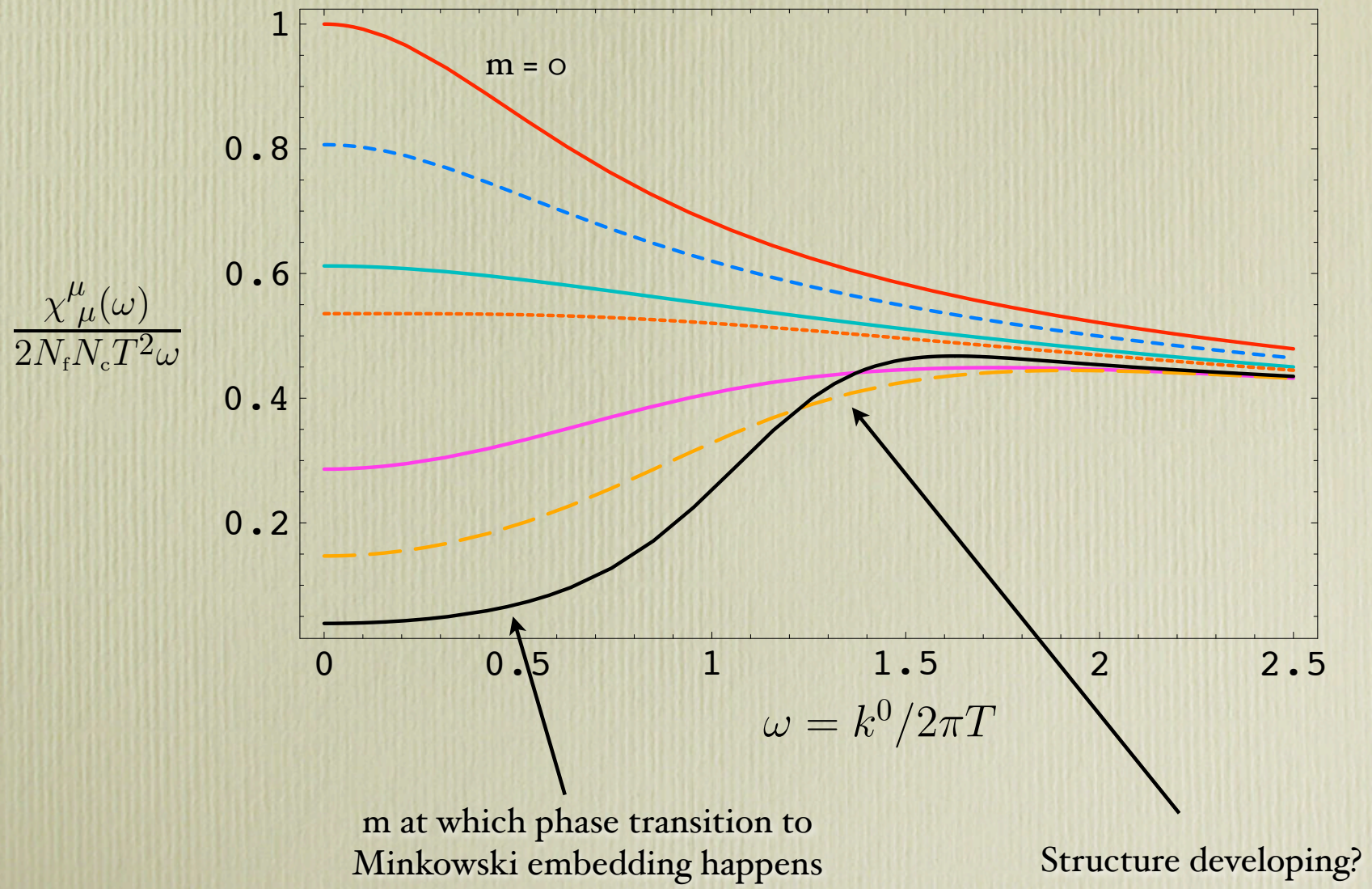
- No obvious comparison of M_q -dependence to pQCD:

Arnold, Moore & Yaffe '01

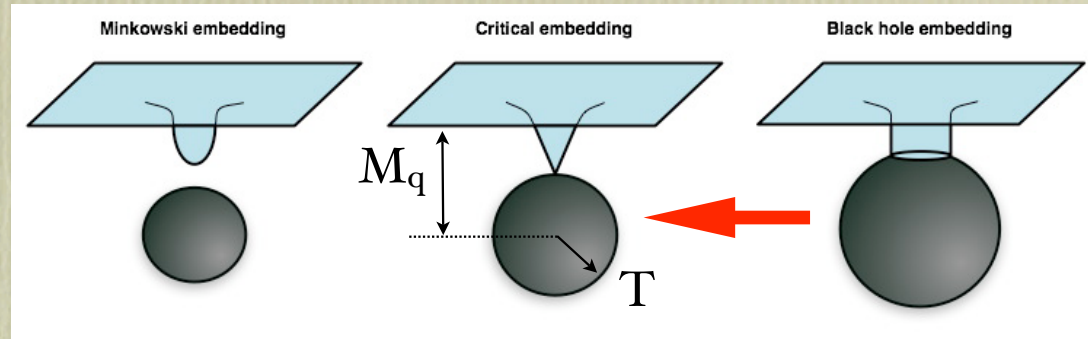
$$M_{\text{thermal}} \sim \sqrt{\lambda} T \gg M_q$$

But this assumes existence of quasi-particles!

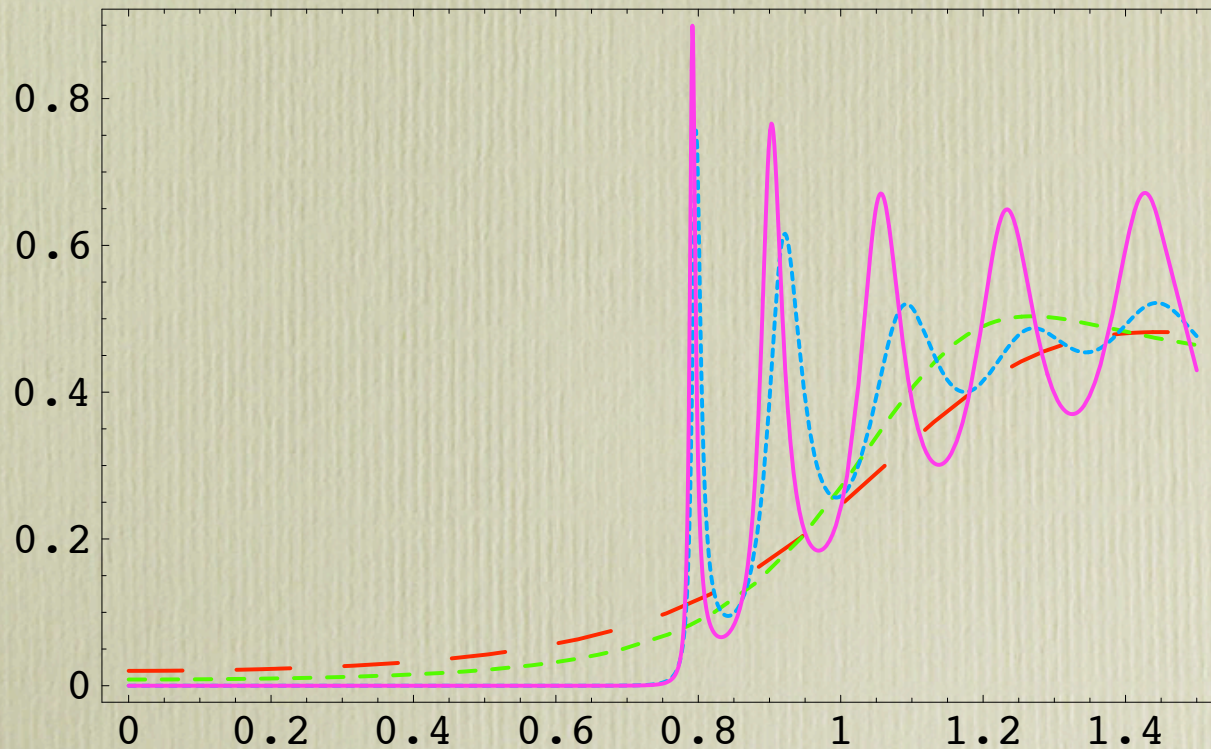
Spectral function for constant $m \sim \frac{M_{\text{mes}}}{T} \sim \frac{M_q}{\sqrt{\lambda}T}$



Approaching the critical embedding:



$$\frac{\chi_{\mu}^{\mu}(\omega)}{2N_f N_c T^2 \omega}$$

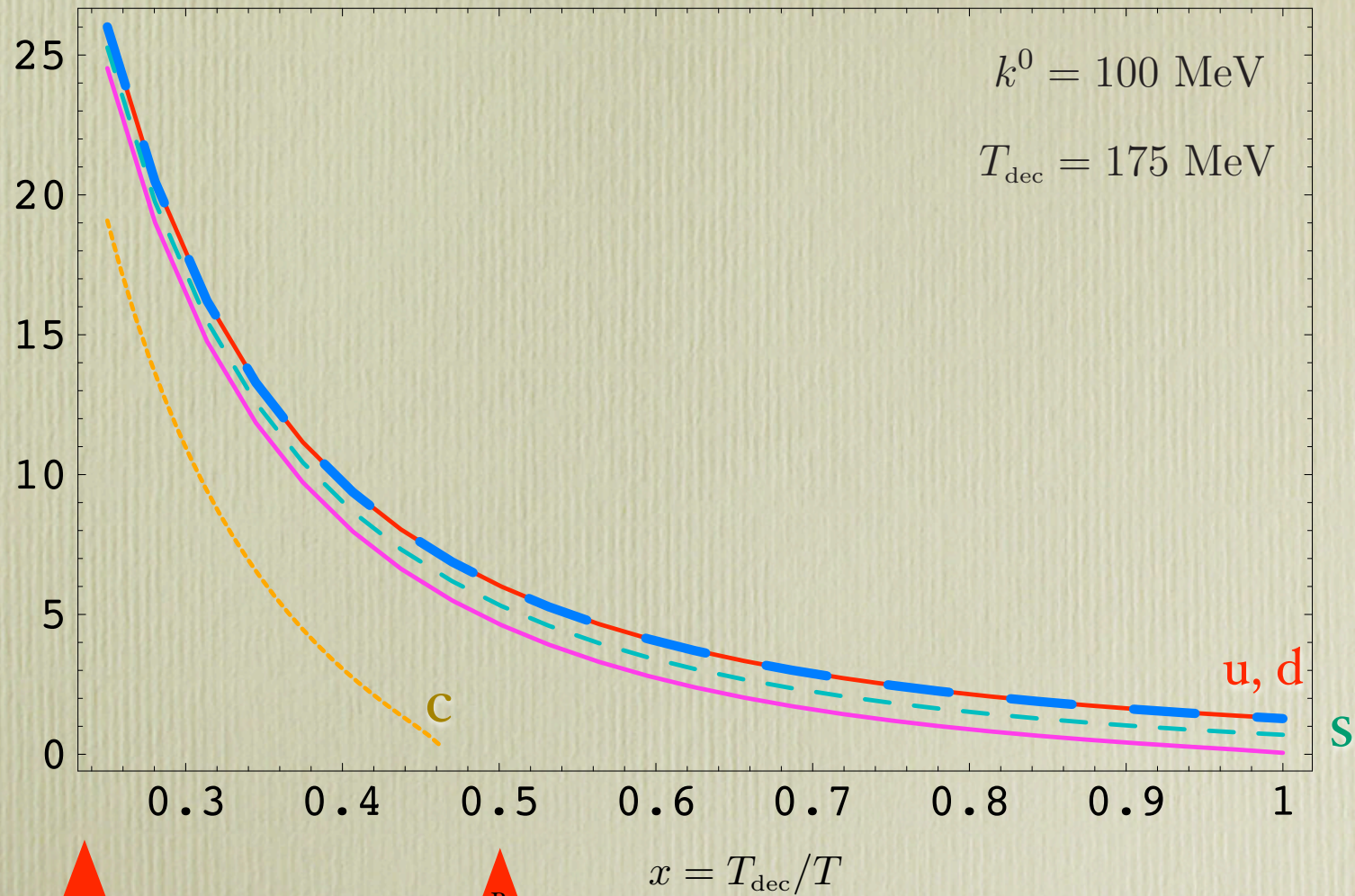


$$\omega = k^0 / 2\pi T$$

To compare with experiment

- Calculate photon production rate:

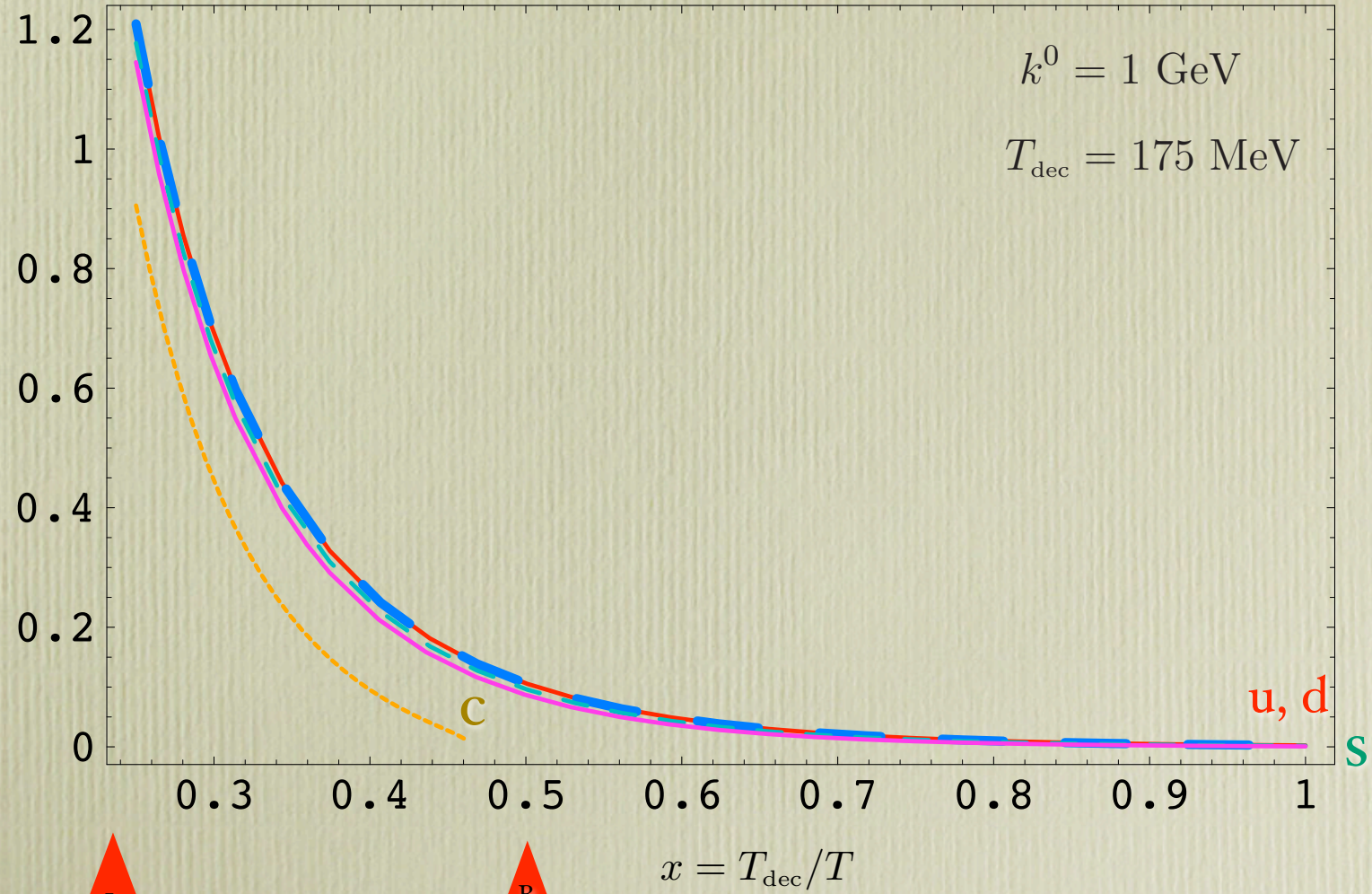
$$\frac{\pi^2}{\alpha_{\text{EM}} N_f N_c (k^0)^2 T_{\text{dec}}} \frac{d\Gamma}{dk^0}$$



L
H
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$$\frac{\pi^2}{\alpha_{\text{EM}} N_f N_c (k^0)^2 T_{\text{dec}}} \frac{d\Gamma}{dk^0}$$



L
H
C

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C

To compare with experiment

- Plug into hydrodynamic simulation of spacetime evolution of the plasma.
- Experimentally distinguish different sources:
QGP photons, prompt photons, decay photons, etc.

v) Finite baryon density.

Kobayashi, D.M., Matsuura, Myers & Thomson '06

Gauge theory

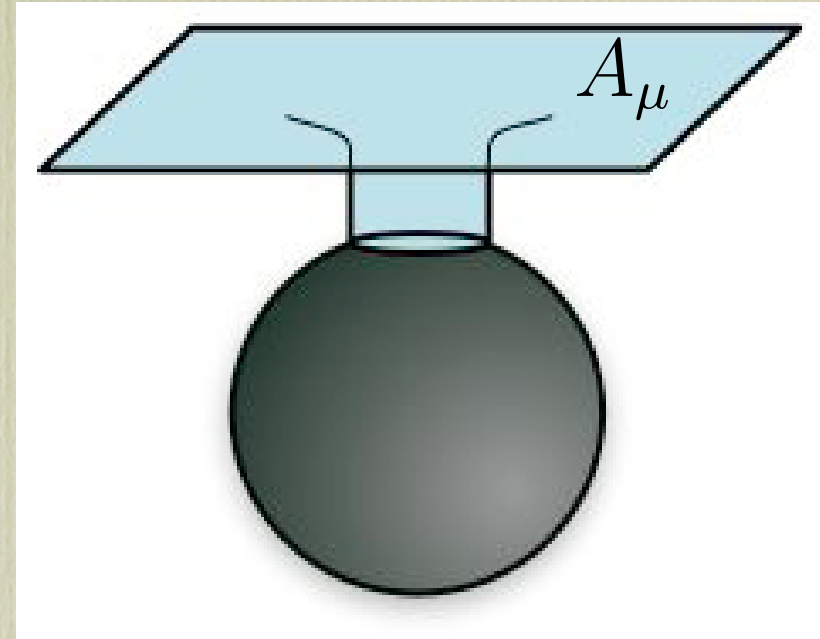
$$U(N_f) \simeq SU(N_f) \times U(1)_B$$

Conserved J_μ

$$H \rightarrow H + \mu_B J^0$$

$$\langle J^0 \rangle = n_B$$

String theory



$$A_0 = \mu_B + \frac{n_B}{r^2} + \dots$$

Gauge theory

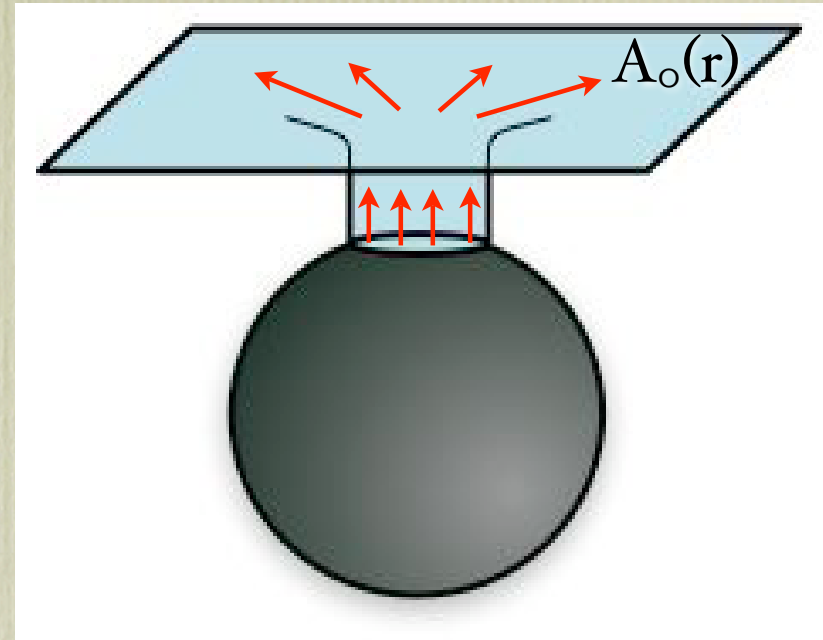
$$U(N_f) \simeq SU(N_f) \times U(1)_B$$

Conserved J_μ

$$H \rightarrow H + \mu_B J^0$$

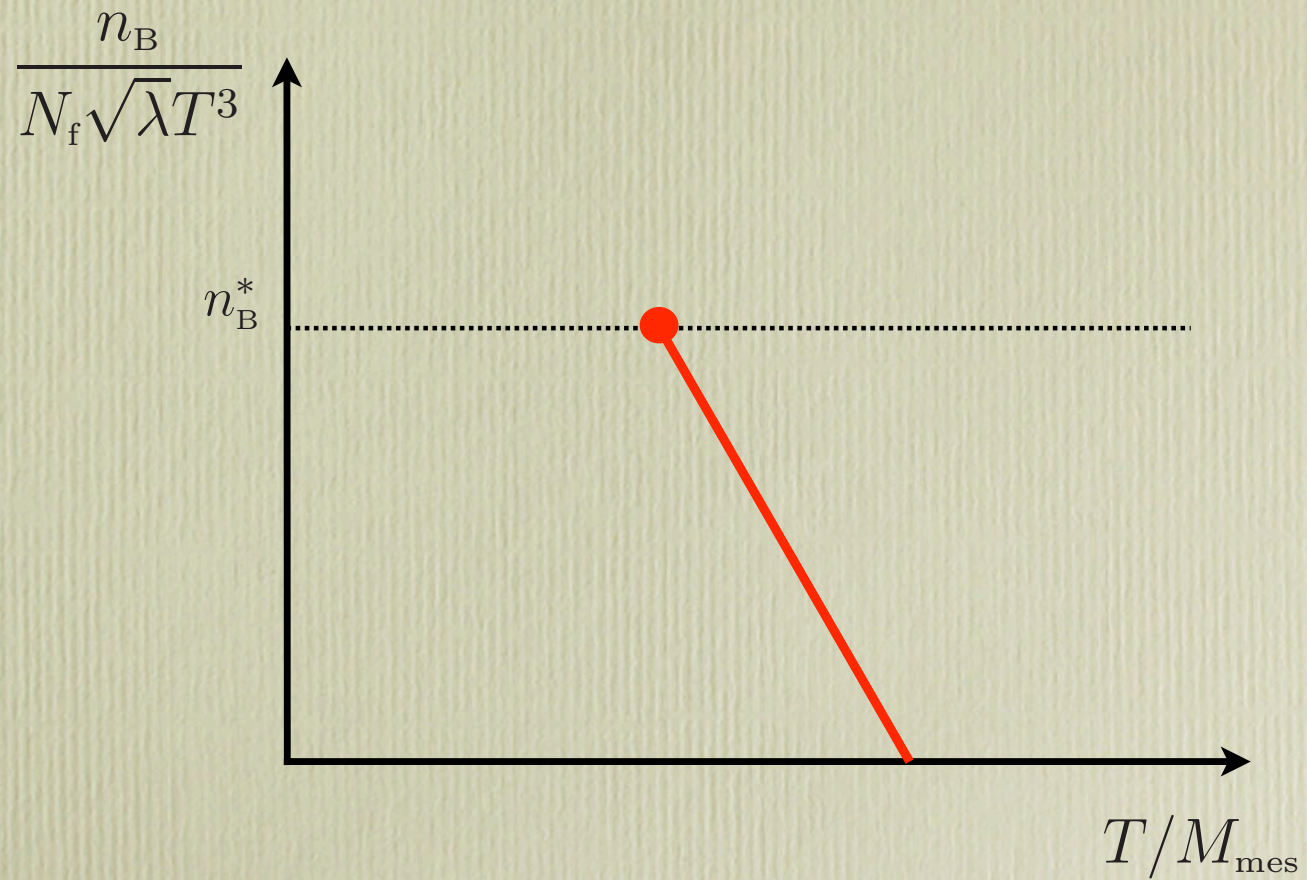
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String theory



$$A_0 = \mu_B + \frac{n_B}{r^2} + \dots$$

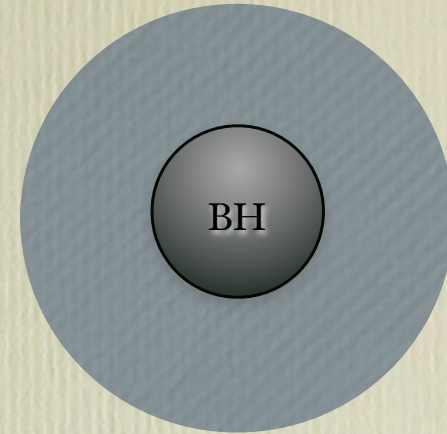
Phase Diagram



vi) Future prospects.

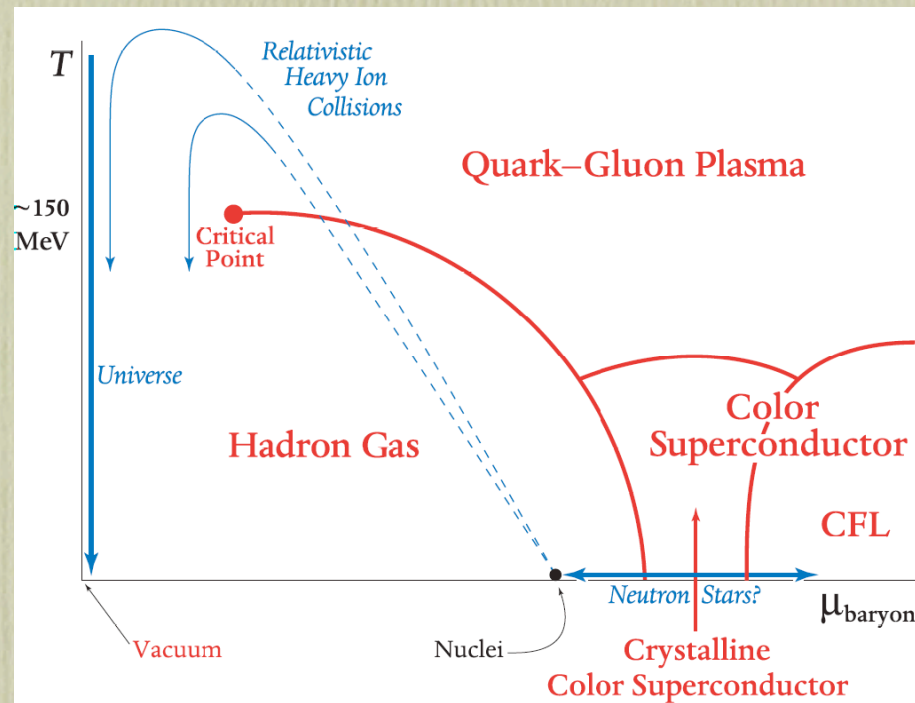
Towards far from equilibrium

Horizons encode properties of QGPs:



- Static \rightarrow Thermodynamics: $S=A/4G$.
- Small perturbations \rightarrow Near equilibrium,
eg. transport coefficients.
- Large perturbations \rightarrow Far from equilibrium,
eg. collective instabilities.

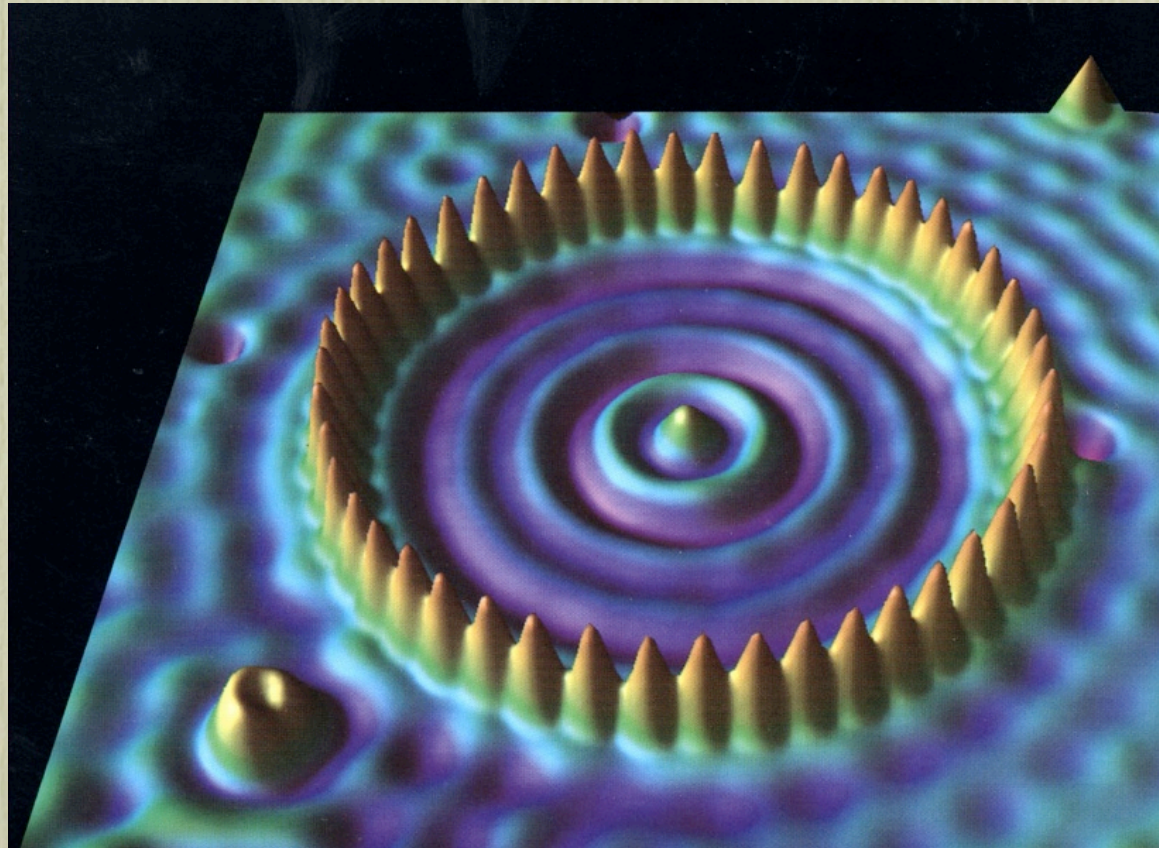
Lots to do at finite density



Caveats: Scalar fields (but not always) and large N_c .

Sakai & Sugimoto

Towards holographic condensed matter



Herzog, Kovtun, Sachdev & Son '07

Hartnoll & Kovtun '07

Hartnoll, Kovtun, Mueller & Sachdev '07

Heavy ion collisions at LHC

$$T_{\text{RHIC}} \sim 2T_{\text{dec}}, \quad T_{\text{LHC}} \sim 4T_{\text{dec}}$$



Thank you.