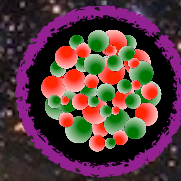
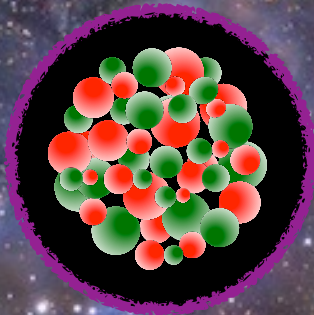


# Resolving Black Holes

via

## *Microstate Geometries*



**Nick Warner, *Strings* '14, June 24, 2014**

Based on Collaborations with:

I. Bena, G. Gibbons, M. Shigemori

arXiv:1305.0957, arXiv:1311.4538 and arXiv:1406.4506

⇒ Two distinct core ideas from *microstate geometries*

1) A string theory *mechanism* to support structure at the horizon scale

⇒ **Two** (at least) **new scales for black-hole physics**

What does horizon-scale microstate structure look like?

Fuzz/Fire/hybrid... other?

2) A semi-classical description of black-hole microstates?

Arising from fluctuations/moduli of microstate geometries

(in the same regime of parameters in which there is an *actual black hole*)

The *superstratum* (BPS):  $S_{semi\ class.} \sim \sqrt{N_1 N_5 N_P}$

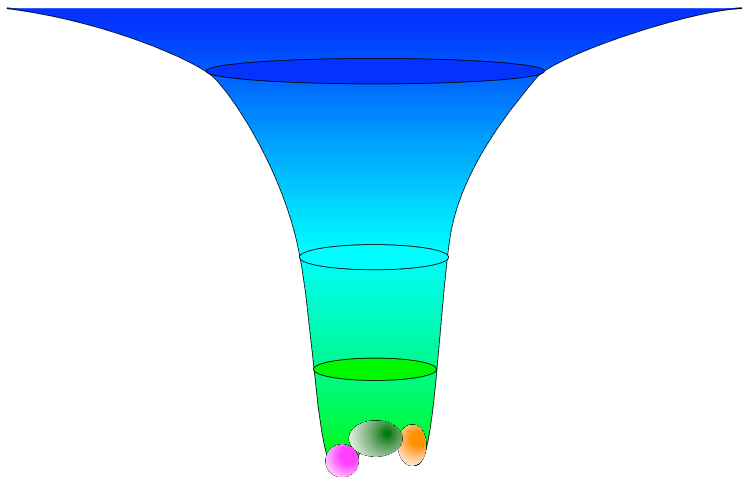
# Fixing the information problem: An old conceit

Recover information (and pure quantum state) from small corrections to GR over the evaporation time scale...

e.g. via stringy or quantum gravity ( $(Riemann)^n$ ) corrections to radiation?

Mathur (2009): **No!** Corrections cannot be small for information recovery  
There must be  $O(1)$  changes to the physics at the horizon scale

## Microstate geometries



A *mechanism* for resolving the *problem* in string theory

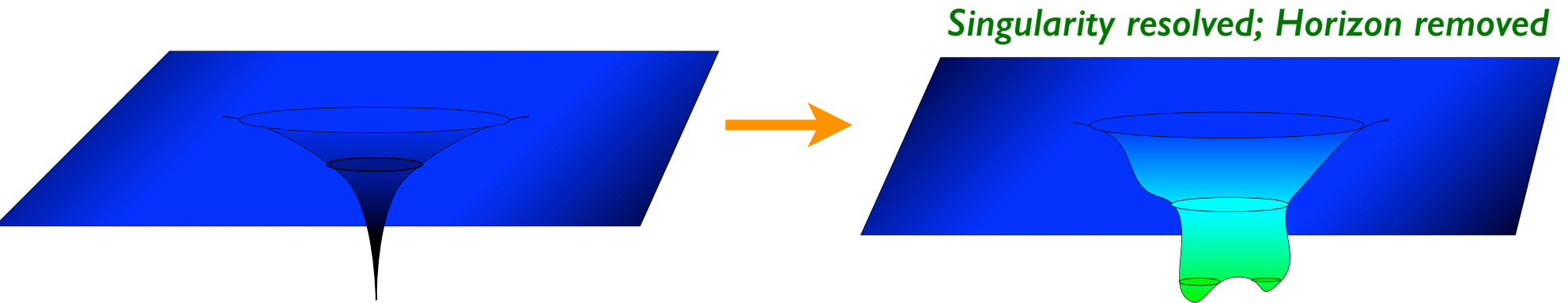
## Firewalls



Unsupported superstructure

# Microstate Geometry Program

*Microstate Geometry*  $\equiv$  **Smooth, horizonless solutions** to the **bosonic** sector of **supergravity** with the same asymptotic structure as a given black hole/ring



**Supergravity** because we seek stringy resolutions on horizon scale

- ▶ **Very long-range effects**  $\Rightarrow$  Massless limit of strings ...
- ▶ Framework within which we can **actually do calculations**

**What is the form of generic, (non-)BPS, time-independent horizonless, smooth solutions in supergravity?**

**Microstate Geometries/solitons** long believed **impossible** because **only the presence of a horizon** can restrict massless fields to a classical lump ...

**Microstate Geometries** exist (how?) ... and lead to new physical issues

- **New physics/scales will emerge from the resolution**
- **What can supergravity tell us about details of microstate structure?**

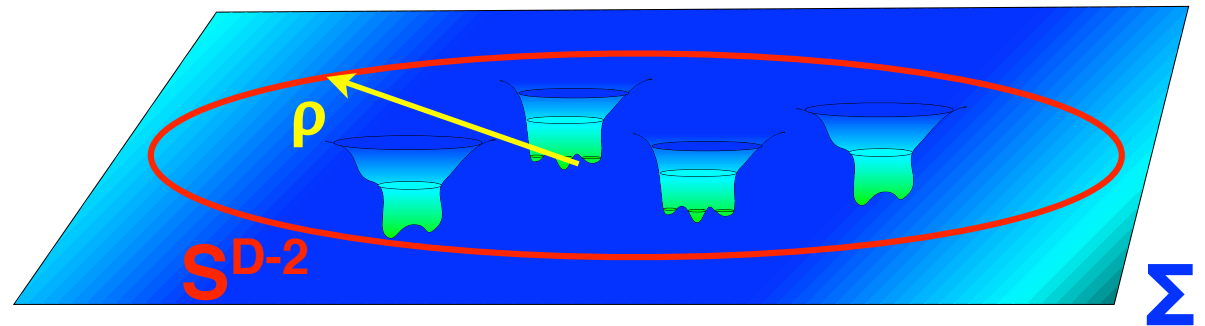
# The Komar Mass/Smarr Formula

If there is time-translation invariance then energy is conserved:

There is a vector field (Killing vector)  $\mathbf{K}$  generating time translations.

$$K^\mu \frac{\partial}{\partial x^\mu} = \frac{\partial}{\partial t}$$

D-dimensional space-time,  
sectioned by hypersurfaces,  $\Sigma$ ,  
with Gaussian (D-2)-spheres,  
 $S^{D-2}$ , at infinity



There is then an associated conserved *ADM mass*:

$$K^\mu K^\nu g_{\mu\nu} = g_{00} \approx -1 + \frac{16\pi G_D}{(D-2) A_{D-2}} \frac{M}{\rho^{D-3}} + \dots \quad \text{at infinity}$$

$$\Rightarrow M = -\frac{1}{16\pi G_D} \frac{(D-2)}{(D-3)} \int_{S^{D-2}} *dK$$

If  $\Sigma$  is smooth with *no interior boundaries*:  $d * dK = -2 * (K^\mu R_{\mu\nu} dx^\nu)$

$$M = \frac{1}{8\pi G_D} \frac{(D-2)}{(D-3)} \int_{\Sigma} *_D (K^\mu R_{\mu\nu} dx^\nu)$$

# Bosonic sector of a generic *massless* (ungauged) supergravity

- Graviton,  $g_{\mu\nu}$
- Scalars,  $\Phi^A$
- Tensor gauge fields,  $F_{(p)}^K$

Bianchi:  $d(F_{(p)}^K) = 0$

Define:  $G_{J,(D-p)} \equiv * (Q_{JK}(\Phi) F_{(p)}^K + \text{Chern Simons terms})$

$Q_{JK}(\Phi)$  = Scalar matrix in kinetic terms

Equations of motion:  $d(G_{J,(D-p)}) = 0$

Assume time-independent matter:

$$\mathcal{L}_K F^I = 0, \quad \mathcal{L}_K \Phi^A = 0 \quad \Rightarrow \quad \mathcal{L}_K G_I = 0$$

Cartan formula for forms:  $\cancel{\mathcal{L}_K} \omega^0 = d(i_K(\omega)) + i_K(\cancel{d}\omega^0)$

$$d(i_K(F_{(p)}^I)) = 0, \quad d(i_K(G_{J,(D-p)})) = 0$$

Define *harmonic* forms,  $H$ :

$$i_K(F_{(p)}^I) = H_{(p-1)}^I + \text{exact} \quad i_K(G_{J,(D-p)}) = H_{J(D-p-1)} + \text{exact}$$

# No Solitons without Topology

Smooth spatial sections with *no interior boundaries*

$$M = \frac{1}{8\pi G_D} \frac{(D-2)}{(D-3)} \int_{\Sigma} *D (K^{\mu} R_{\mu\nu} dx^{\nu})$$

Equations of motion imply

$$M = \text{const.} \int_{\Sigma} \left[ H_{I(D-p-1)} \wedge F_{(p)}^I + H_{(p-1)}^J \wedge G_{J(D-p)} \right]$$

Gibbons + NPW 1305.0957; Haas 1405.3708

- Mass *can be* topologically supported by the cohomology  $H^*(\Sigma, \mathbb{R})$

Stationary end-state of star held up by topological flux ...

- A new object: A *Topological Star*
- *Black-Hole Microstate?*

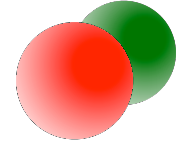
- No spatial topology  $\Rightarrow M = 0 \Rightarrow$  Space-time is flat/empty

Only assumed time independence: **Not simply for BPS objects**

*Applies to all time-independent smooth remnants in massless ungedged supergravity*

# A Decade of *BPS* Microstate Geometries

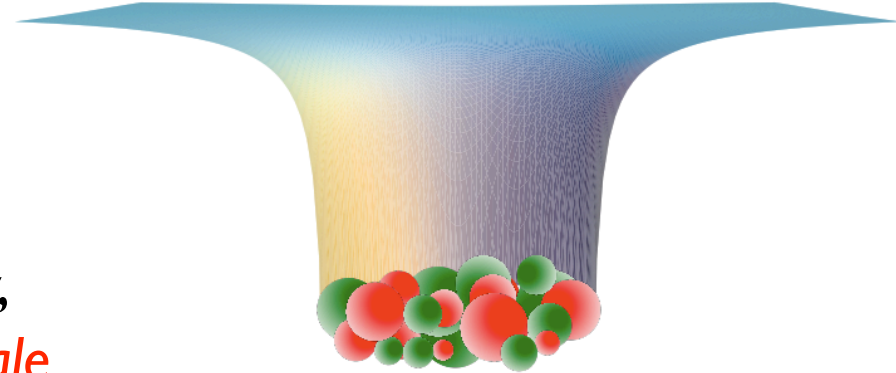
Bubbled geometries in *five* or *six* dimensions  $\Rightarrow$  *2* or *3* cycles



★ There are vast families of smooth, horizonless microstate geometries

★ New physics *at the horizon scale*

$\Rightarrow$  The cap-off and the non-trivial topology, “bubbles,” arise at the original *horizon scale*



★ Families of solutions: Large moduli spaces of cycles; fluctuations around cycles

★ Special class: KK reduction yields multi-centered solutions of Denef

★ There are *scaling microstate geometries* with *AdS throats* that can be made *arbitrarily long* but cap off smoothly

★ Holography in the long AdS throat:

*All these solutions represent black-hole microstates*

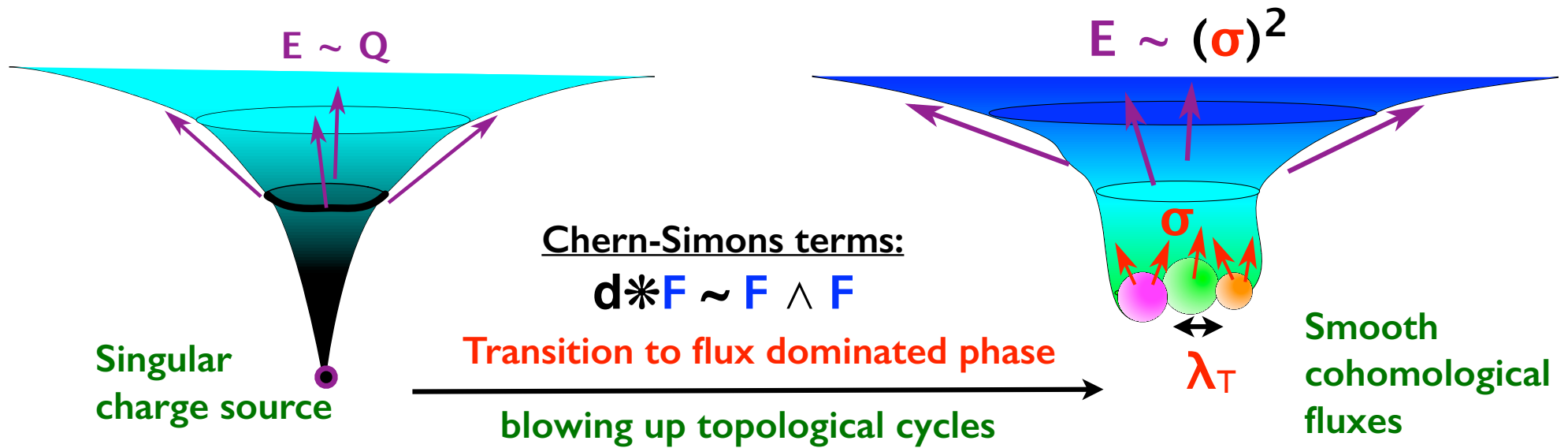
$\Rightarrow$  Semi-classical sampling of black-hole microstate structure

*“Topological stars” = coherent microstates of black holes*

★ New physical scales ...



# Scale 1: The Order Parameter of the Geometric Phase



This is an example of a phase/geometric transition in string theory ...

*Analogous to holography of confinement and chiral symmetry breaking.*

- ★ Magnitude of fluxes,  $\sigma$  = Order parameter of new phases
- ★ Size of the bubbles,  $\lambda_T$  = *Transition Scale* is a new scale in the topological phase

Supergravity equations  $\Rightarrow \lambda_T \sim$  Magnitude of fluxes,  $\sigma$

Balance: Gravity  $\leftrightarrow$  Flux expansion force

Classically: *Freely choosable parameter*. Can have  $\lambda_T \gg \ell_p$

Quantum mechanics: Could  $\lambda_T$  be dynamically generated?

Black holes: Could large  $\lambda_T$  be entropically favored?

## Scale 2: The Energy Gap

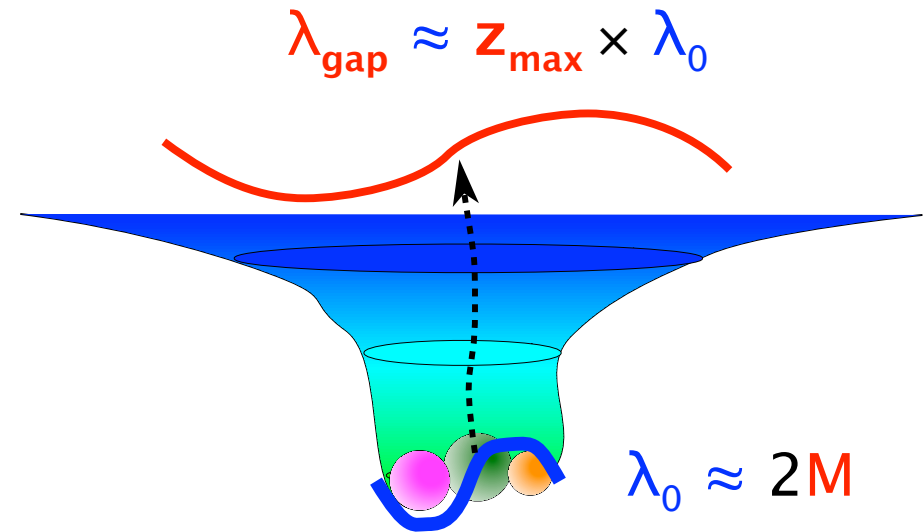
$\lambda_{\text{gap}}$  = maximally redshifted wavelength, at infinity of lowest collective mode of bubbles at the bottom of the throat.

$$E_{\text{gap}} \sim (\lambda_{\text{gap}})^{-1}$$

- ★ The gap is determined by “maximum redshift,”  $z_{\text{max}}$ , and size of black hole

Traditional black holes:  $E_{\text{gap}} = 0$

- ★  $E_{\text{gap}}$  determines where microstate geometries begin to differ from black holes



### BPS black holes

Semi-classical quantization of the moduli of the geometry:

- ★ The throat depth, or  $z_{\text{max}}$ , is *not* a free parameter
- ★  $E_{\text{gap}}$  is determined by the flux structure of the geometry

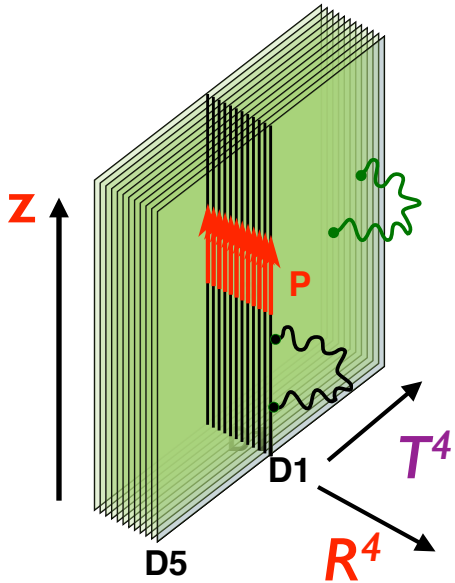
*Exactly matches*  $E_{\text{gap}}$  for the stringy excitations underlying the original state counting of Strominger and Vafa .....

Bena, Wang and Warner, arXiv:0706.3786

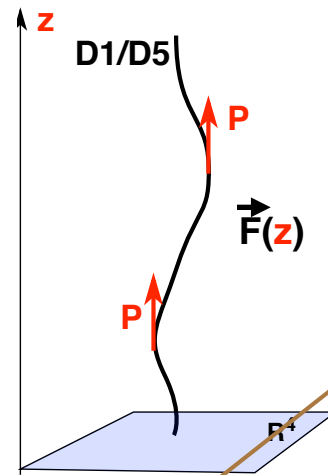
de Boer, El-Showk, Messamah, Van den Bleeken, arXiv:0807.4556 arXiv:0906.0011

# Semi-classical Microstate Structure: *Superstrata* on $R^{5,1} \times T^4$

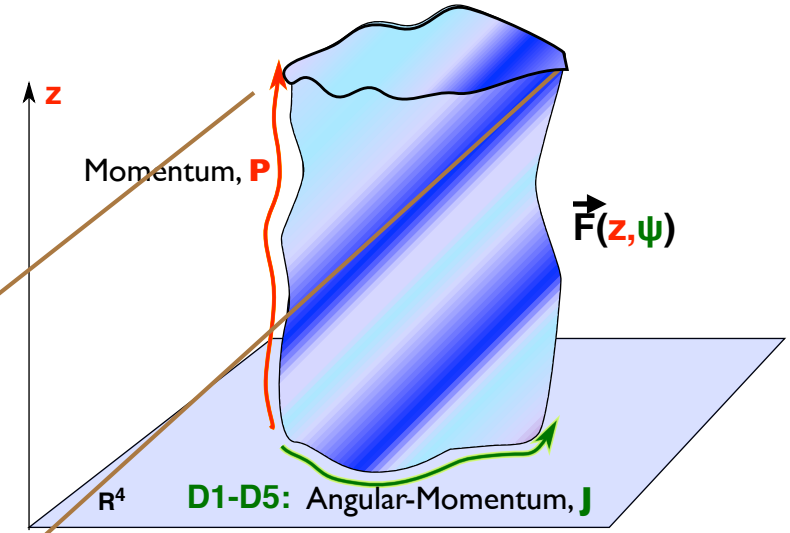
IIB: D1-D5-P system compactified on  $T^4$  (or  $K3$ )



Six Dimensions: Profile in  $R^{5,1}$

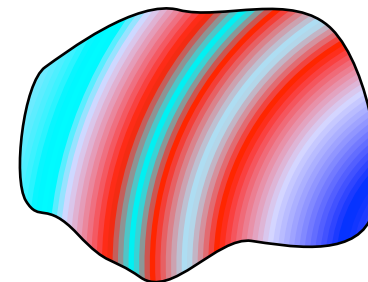
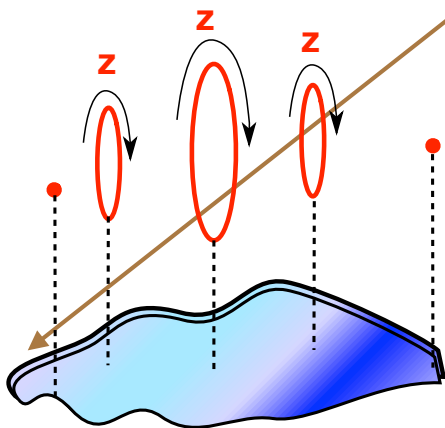


Add KKM dipole and Angular Momentum

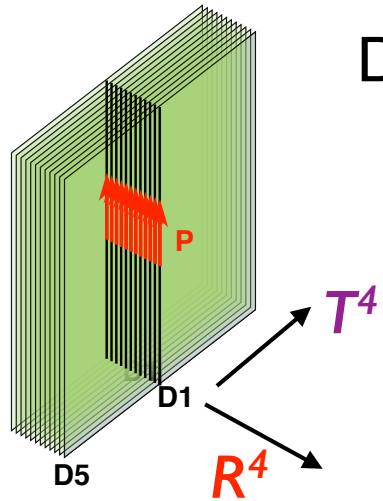


*Smooth BPS* configurations that depend upon *functions of two variables*:  $\vec{F}(z, \psi)$   
 Back-reacted geometry: 3 homology

$\Rightarrow$  BPS shape modes on 3 cycles  
*functions of two variables*



# D1-D5-P System: Microstate Counting *Strominger-Vafa*



D1-D5 SCFT Fields  $X_{(r)}^{\dot{A}A}(z, \bar{z})$   $\psi_{(r)}^{\alpha\dot{A}}(z)$   $\tilde{\psi}_{(r)}^{\dot{\alpha}\dot{A}}(\bar{z})$   
 $r = 1, \dots, N = N_1 N_5$

$$\frac{(T^4)^N}{S_N} \quad c = 6N = 6N_1 N_5 \quad (4,4) \text{ supersymmetry}$$

$(A, \dot{A})$  = spinor indices on the  $T^4$   
 $(\alpha, \dot{\alpha})$  = spinor indices on  $R^4$  transverse to branes

**R-symmetry** = Rotations in  $R^4$  transverse to branes  
 =  $SO(4) = SU(2)_L \times SU(2)_R$

Define:  $J_{(r)}^{\alpha\beta}(z) \equiv \frac{1}{2} \psi_{(r)}^{\alpha\dot{A}}(z) \epsilon_{\dot{A}\dot{B}} \psi_{(r)}^{\beta\dot{B}}(z)$ ,  $\tilde{J}_{(r)}^{\dot{\alpha}\dot{\beta}}(\bar{z}) \equiv \frac{1}{2} \tilde{\psi}_{(r)}^{\dot{\alpha}\dot{A}}(\bar{z}) \epsilon_{\dot{A}\dot{B}} \tilde{\psi}_{(r)}^{\dot{\beta}\dot{B}}(\bar{z})$

= **CFT Degrees of freedom “visible” in  $R^4$**

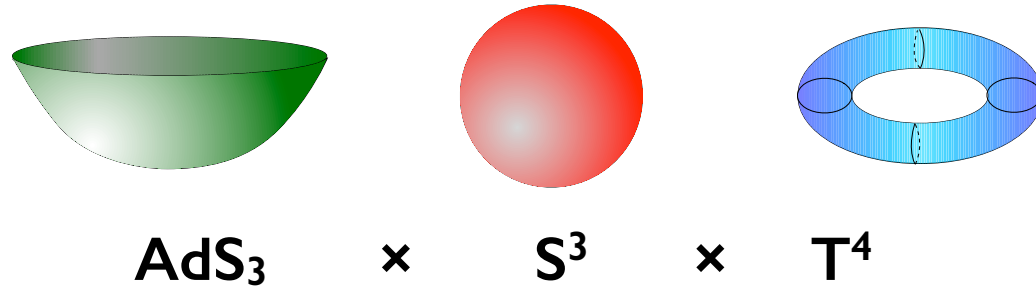
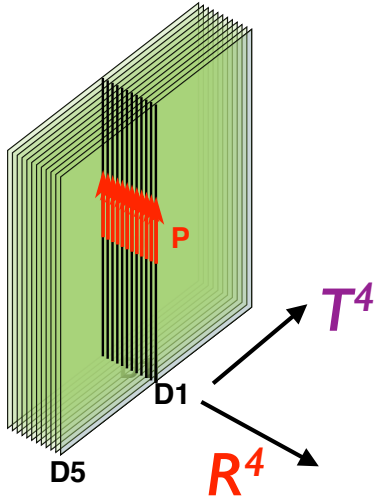
$$= \frac{(SU(2)_{(1),L} \times SU(2)_{(1),R})^N}{S_N} \quad c = N = N_1 N_5$$

$\frac{1}{4}$  BPS states = (R,R)-ground states

$\frac{1}{8}$  BPS states = (any left-moving state, R ground state) **Momentum,  $P = L_{0,\text{left}}$**

The left-handed currents,  $J_{(r)}^{\alpha\beta}(z)$ , ( $c = N_1 N_5$ ) create left-moving momentum states visible in  $R^4 \Leftrightarrow$  **BPS Shape modes** of the *superstratum*/ $S^3$

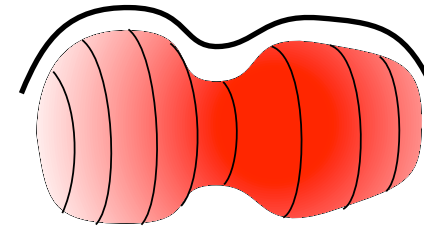
# The Holographic Dual: $AdS_3 \times S^3 \times T^4$



$AdS_3 \times S^3 \times T^4$   
 Modes:  $SU(2)_L \times SU(2)_R$  quantum numbers  $(j, m; \hat{j}, \hat{m})$   
 $|j - \hat{j}| = \text{space-time spin of underlying field}$

$\frac{1}{4}$  BPS states = (R,R)-ground states  $\leftrightarrow$  quantum numbers  $(j, j; \hat{j}, \hat{j})$

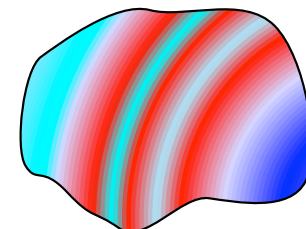
- $\rightarrow$  D1-D5 supertube shape modes on  $S^3$
- $\leftrightarrow$  *one arbitrary function, Fourier modes,  $j$*   
 Lunin and Mathur; Lunin, Maldacena and Maoz; Mathur; Skenderis and Taylor ...



$\Rightarrow$  Semi-classical entropy of a supertube:  $S \sim \sqrt{Q_1 Q_2} \sim Q$

$\frac{1}{8}$  BPS states: Act with modes of  $J_{(r)}^{\alpha\beta}(z) \leftrightarrow$  quantum numbers  $(j, m; \hat{j}, \hat{j})$

- $\rightarrow$  D1-D5 *superstratum* shape modes on  $S^3$
- $\leftrightarrow$  *two arbitrary functions, Fourier modes,  $(j, m)$*

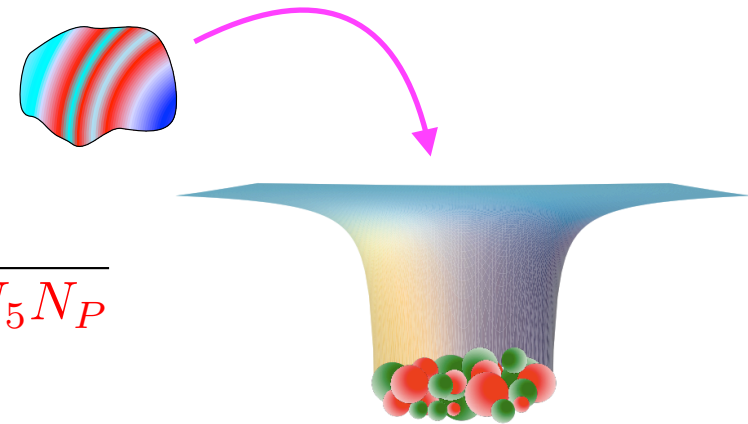


# Semi-classical black-hole microstate structure

★ Linearized supergravity modes can *at least* capture, semi-classically, the momentum excitations corresponding to the CFT currents,  $J_{(r)}^{\alpha\beta}(z)$ .

★ Deep, scaling microstate geometries have  $E_{\text{gap}} \sim (N_1 N_5)^{-1}$

Add momentum charge  $N_P$  using  $J_{(r)}^{\alpha\beta}(z)$



$$\Rightarrow S_{\text{semi class.}} = 2\pi \sqrt{\frac{1}{6} N_1 N_5 N_P} \sim \sqrt{N_1 N_5 N_P}$$

★ Semi-classical quantization gives a dense enough sampling of microstate structure to recover entropy corresponding to a macroscopic horizon scale

$\Rightarrow$  *Typical microstates must have the scale of the original black-hole horizon?*

## Summary

- String theory has new phases dominated by topological fluxes that can *prevent the formation of black holes* → *Topological Stars/Black-hole microstates*
- Transition to new phase  $\leftrightarrow$  Formation of bubbles supported by flux  
→ *Order parameter and new scale in Nature:  $\lambda_T = \text{Transition Scale}$*
- The new phase smoothly caps-off the space-time before a horizon forms:  
→ *Limits the red-shift and the lowest-energy states:  $E_{\text{gap}} > 0$*
- The new phases represent new “*infra-red*” *vacua* of string theory  
*This viewpoint is a natural and direct outgrowth of holographic field theories ...*
- Discussion of near-horizon physics, like the infall problem and even firewalls, will be enriched/clarified by separating  $\lambda_T$  and  $E_{\text{gap}}$  from the Planck scale.  
*Ignoring this possibility is probably a serious mistake ...*
- *Vast families of BPS examples explicitly constructed*
- Superstrata/BPS fluctuations as functions of two variables give *semi-classical entropy with correct growth as a function of the charges*
- These ideas can be extended to *non-BPS, extremal* and *near-extremal* ...