

Black holes in AdS5 and phases of 4d SYM

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Does the most general susy index in $\mathcal{N} = 4$ SYM capture the susy BH entropy in $\text{AdS}_5 \times S^5$?

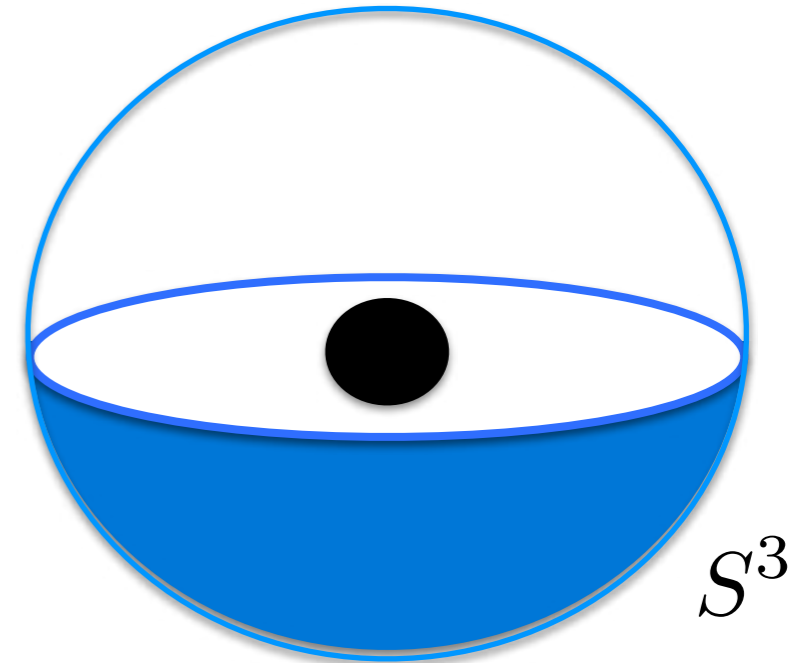


$SU(N)$ $\mathcal{N} = 4$ SYM

$\frac{1}{16}$ -BPS states

time ↑

$$\frac{1}{N^2} = G_N$$



$\text{AdS}_5 (\times S^5)$

$\frac{1}{16}$ -BPS BH

$$\mathcal{I}_N(\tau) = \text{Tr} (-1)^F e^{2\pi i \tau q}$$

$$\log \mathcal{I}_N \stackrel{?}{=} O(N^2)$$

$$S_{\text{BH}} = \frac{A_H}{4G_N} = O(N^2)$$

[Sundborg '99; Aharony, Marsano, Minwalla, Papadodimas, van Raamsdonk '03; Kinney, Maldacena, Minwalla, Raju '05]

[Gutowski, Reall '04; Chong, Cvetič, Lu, Pope '05; Kunduri, Lucietti, Reall '06]

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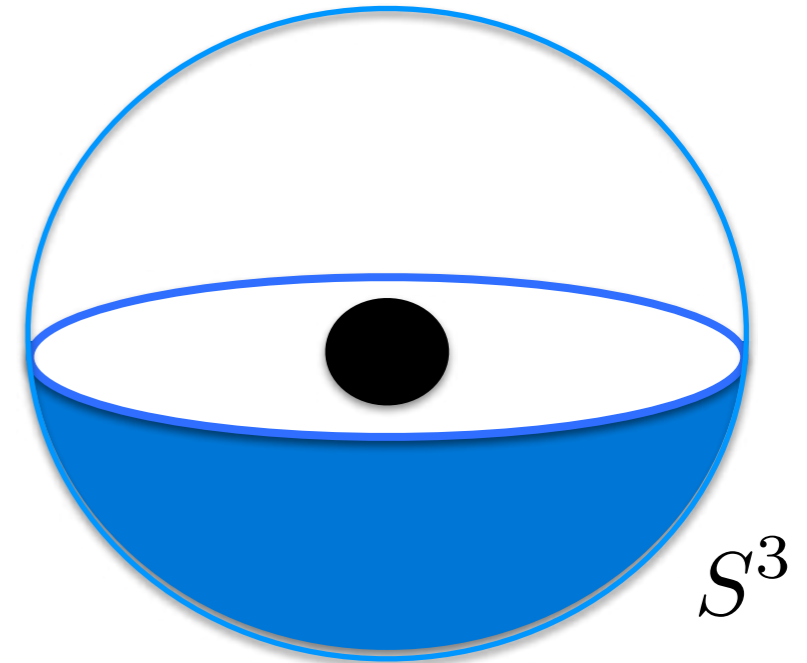


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[New progress: Cabo-Bizet, Cassani, Martelli, S.M. '18; Choi, Kim, Kim, Nahmgoong '18; Benini, Milan '18,]

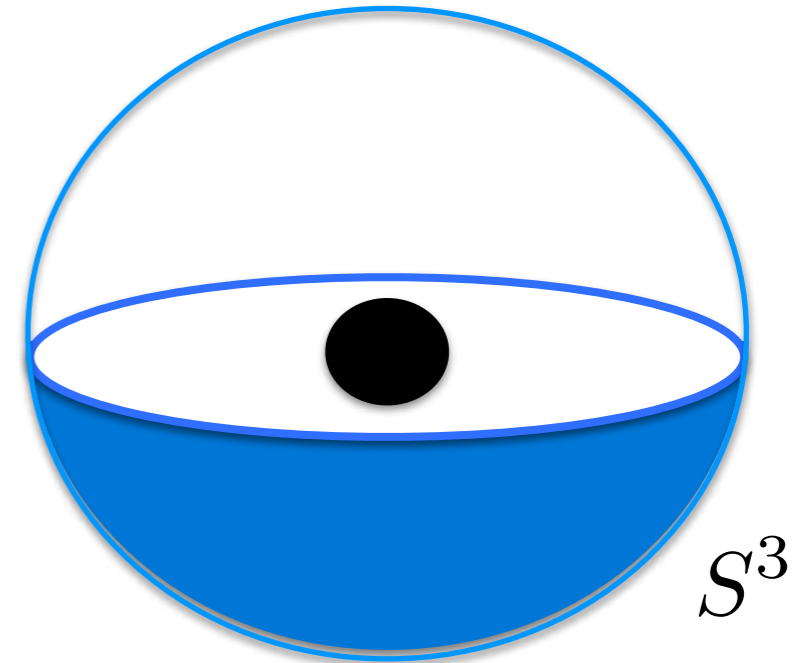
Does the most general index in $\mathcal{N} = 1$ SCFT4 capture the susy BH entropy in dual AdS5 ?



$\mathcal{N} = 1$ SCFT

$\frac{1}{4}$ -BPS states

time ↑



$AdS_5 (\times M_5)$

$\frac{1}{4}$ -BPS BH

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In the simplest setting, sup.conformal index & BHs are labelled by one quantum number

$$\mathcal{N} = 1 \text{ SCFT on } S^3$$

\downarrow

$$U(1)_R \quad Q \quad SO(4)$$

$(2J = J_1 + J_2) \quad (J_1, J_2)$

$$\mathcal{I}_N(\tau) = \text{Tr} (-1)^F e^{2\pi i \tau (2J + Q)}$$

$$= \sum_{n \in \mathbb{Z}} d_N(n) e^{2\pi i \tau n/3}$$

$$S_{\text{BH}} = N^2 \sqrt{3q^2 - 4j}$$

$$= N^2 s(n/N^2)$$

Minimal AdS₅ sugra

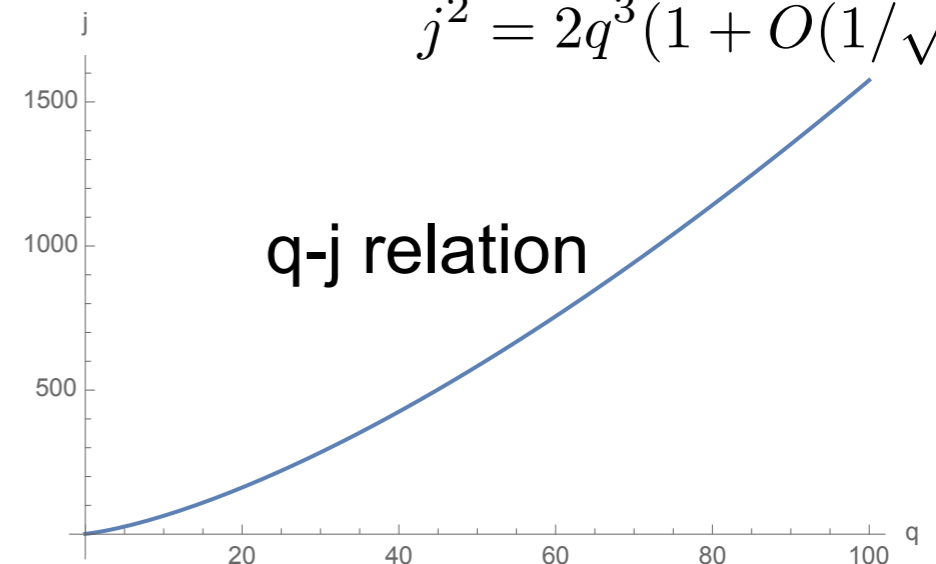
$$g_{\mu\nu}, A_\mu \quad G_N = 1/N^2$$

$\frac{1}{4}$ -BPS BH charges

$$(J_1, J_2, Q) = N^2 (j_1, j_2, q)$$

Simplest: $j_1 = j_2 = j \neq 0$
 [Gutowski, Reall '04]

$$j^2 = 2q^3 (1 + O(1/\sqrt{q}))$$



What is the meaning of BH/exponential growth of states in the index?

$$\log |d_N(n)| \stackrel{?}{\sim} N^2 s(n/N^2)$$

Large- N
limit

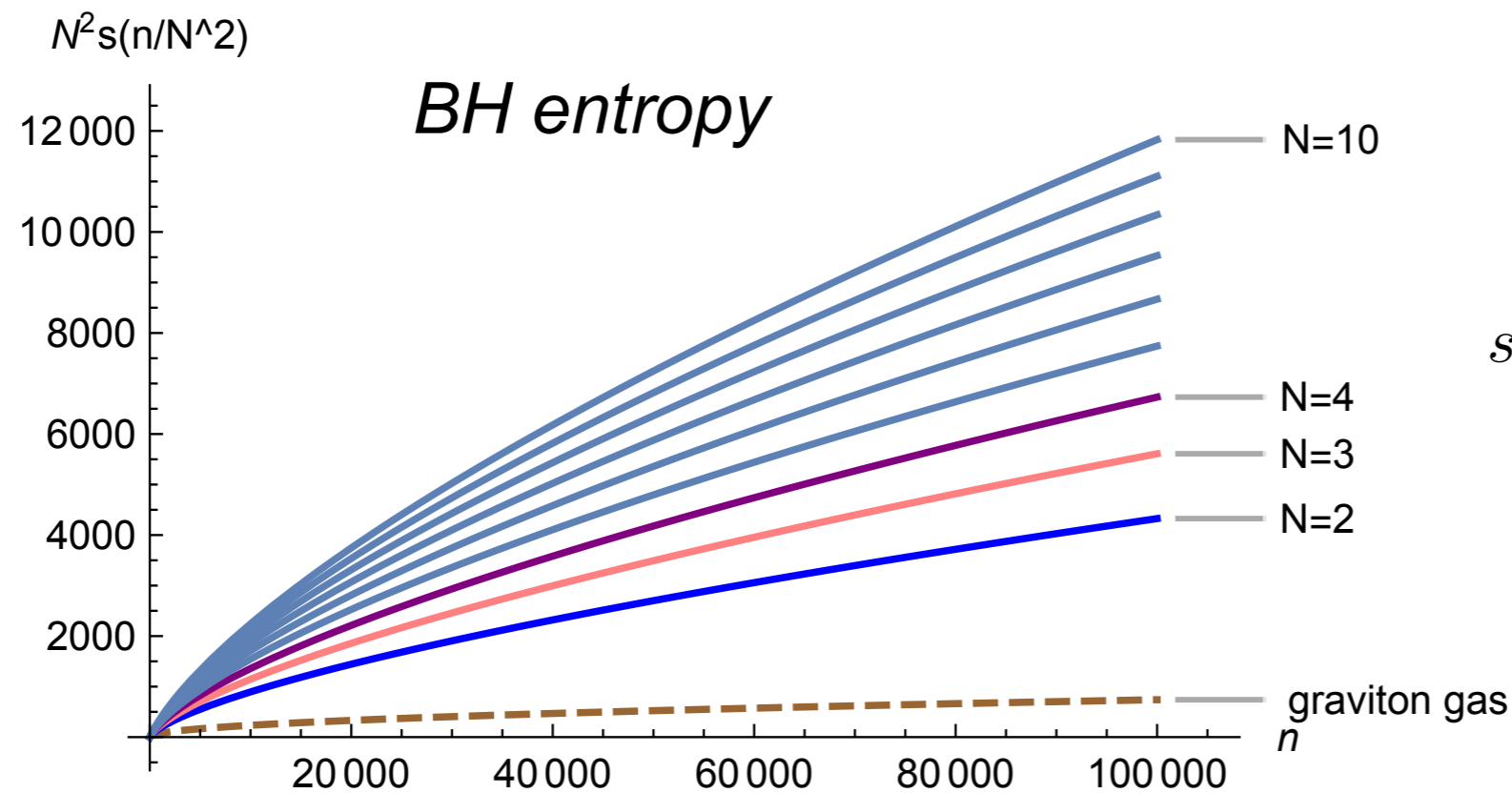
$$n, N \rightarrow \infty$$

$$\text{fixed } \nu = n/N^2$$

Cardy-
like limit

$$n/N^2 \rightarrow \infty$$

$$\text{fixed } N$$



$$s(\nu) \underset{\nu \rightarrow \infty}{\sim} \frac{\pi}{2 \cdot 3^{1/6}} \nu^{2/3} + \dots$$

The N=4 superconformal index is an integral over unitary matrices

[Romelsberger '05;
Kinney, Maldacena, Minwalla, Raju '05]

$$\mathcal{I}_N(\tau) = \int [DU] \exp\left(\sum_{n=1}^{\infty} \frac{1}{n} f(n\tau) \operatorname{tr} U^n \operatorname{tr} (U^\dagger)^n\right)$$

Eigen-values $\{e^{2\pi i u_i}\}_{i=1, \dots, N}$

“Single-letter index”

$$f(x) = \frac{3x^2 - 3x^4 - 2x^3 + 2x^6}{(1-x^3)^2}$$

$$\underline{u} \equiv \{u_i\}_{i=1, \dots, N} \quad \int [D\underline{u}] = \frac{1}{N!} \prod_{i=1}^N \int_0^1 du_i$$

$$\mathcal{I}_N(\tau) = \int [D\underline{u}] \exp\left(-\sum_{i,j=1}^N V_\tau(u_i - u_j)\right)$$

The potential is governed by the *elliptic gamma function*

$$\exp(-V_\tau(z)) = \gamma_e(z + 2\tau; \tau) \gamma_e\left(z + \frac{2}{3}\tau; \tau\right)^3$$

[Dolan-Osborn '08]

$$\gamma_e(z; \tau) = \prod_{n=1}^{\infty} \frac{(1 - q^{n+1} / \zeta)^n}{(1 - q^{n-1} \zeta)^n}$$

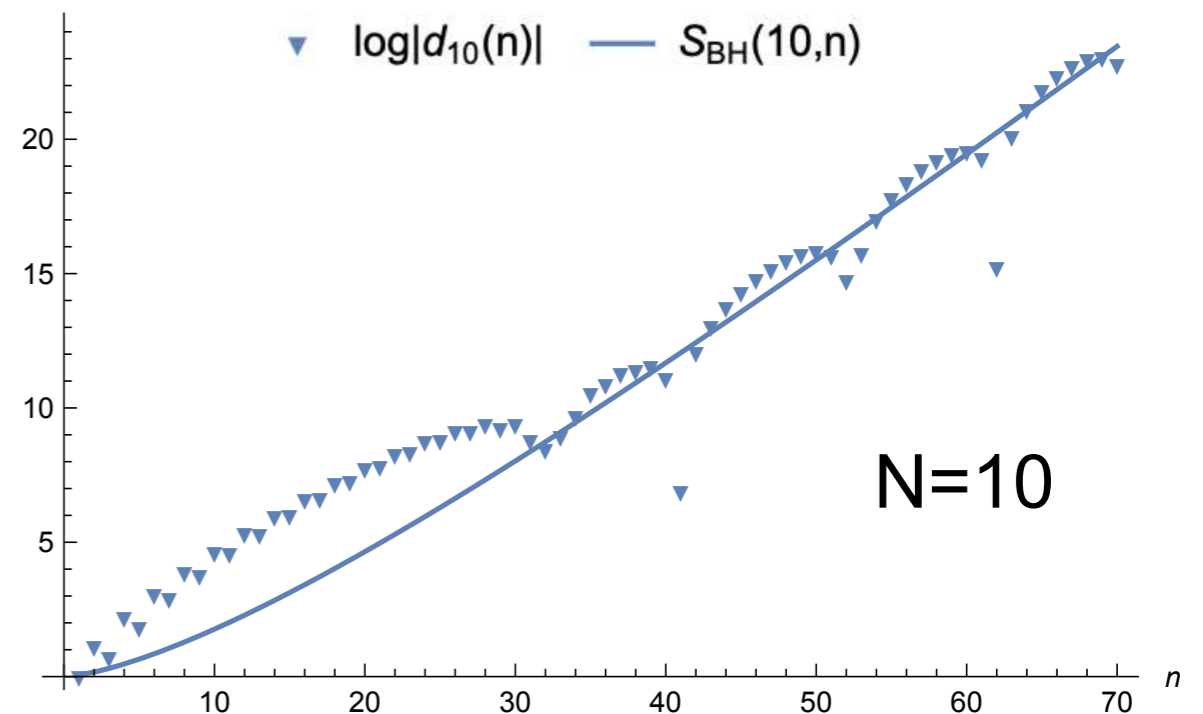
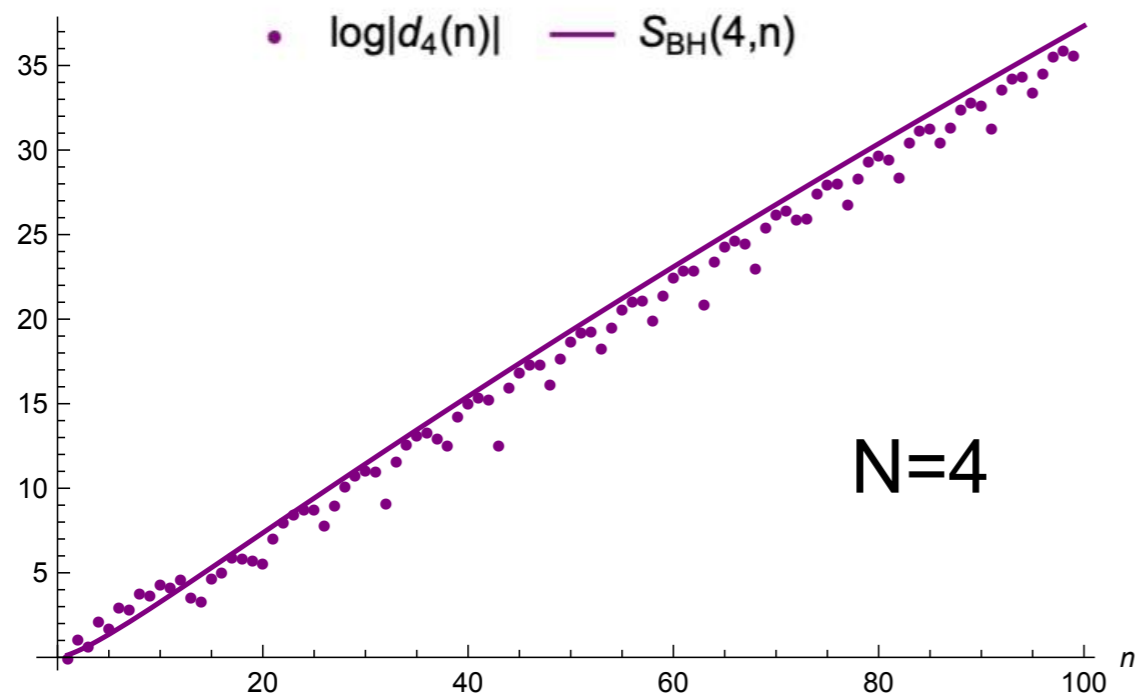
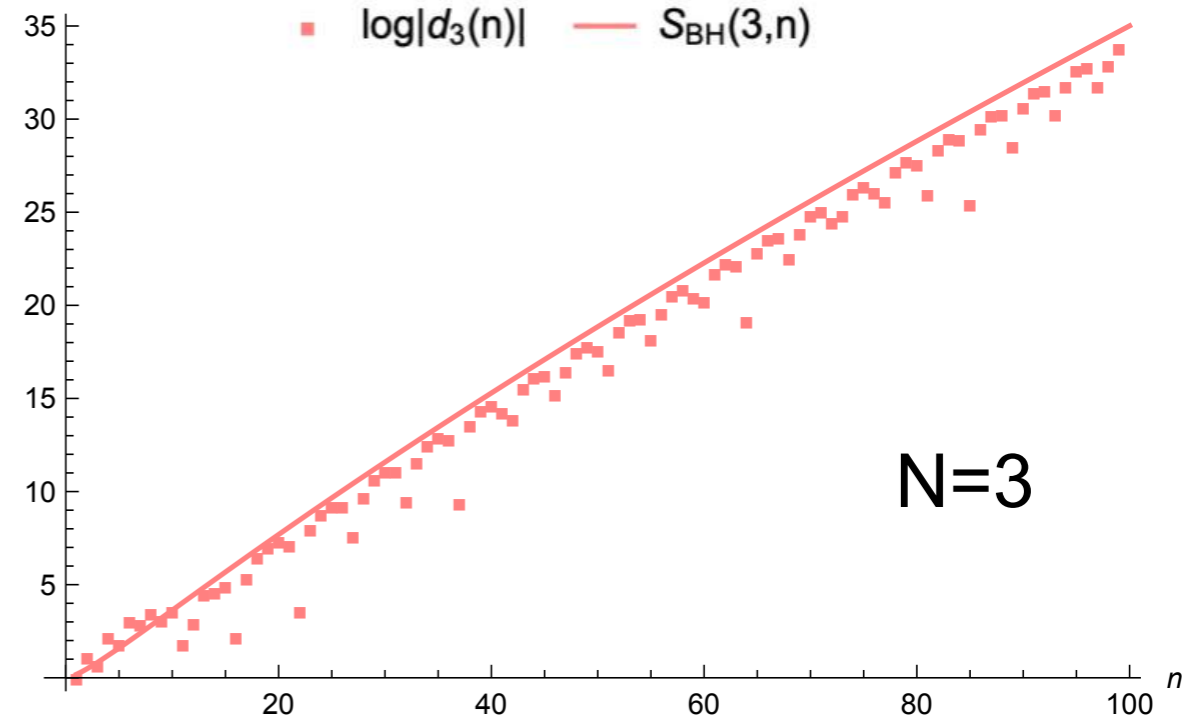
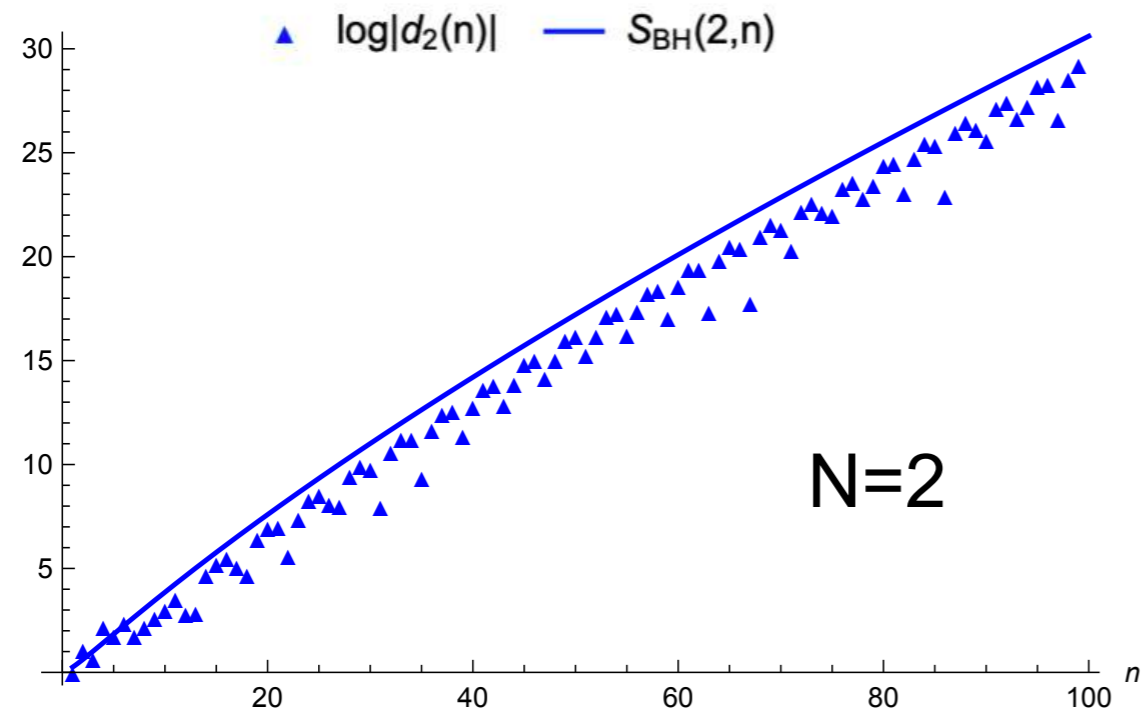
$$\tau \in \mathbb{H}$$

$$q = e^{2\pi i \tau}$$

$$\zeta = e^{2\pi i z}$$

Comparison for N=4 SYM (Cardy-like limit)

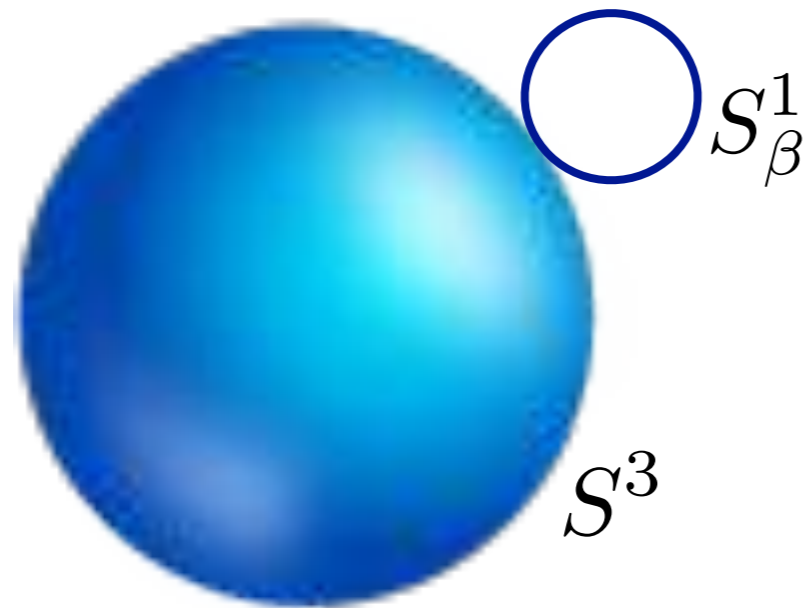
[S.M. '20]



[related numerical work in Agarwal, Choi, Kim, Kim, Nahmgoong '20]

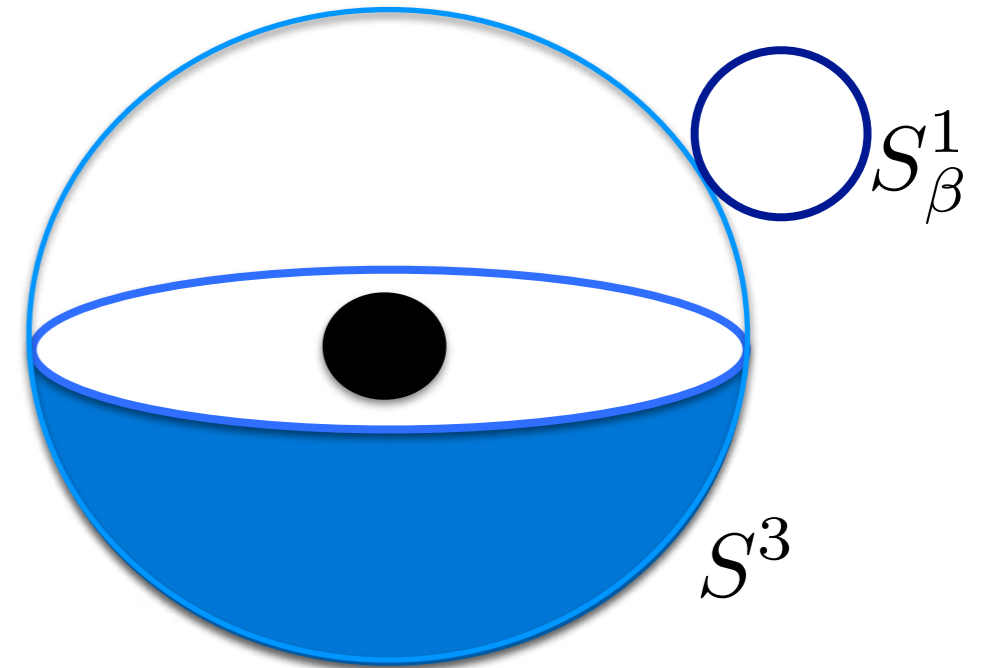
AdS boundary conditions lead to the canonical ensemble \Rightarrow phase structure

[Gibbons-Hawking-Page '80s]



$$\lambda = \ell_{\text{AdS}}^4$$

$$\frac{1}{N^2} = G_N$$



$$Z_{\text{CFT}}(\lambda, N; \beta) = Z_{\text{AdS}}(\ell_{\text{AdS}}, G_N; \beta)$$

$$= \text{Tr} e^{-\beta H} = \sum_{\alpha \in \text{Saddles}} \exp(-\mathcal{F}^\alpha)$$

$$= \sum_n d_{\text{micro}}(n) e^{-\beta E_n} = e^{-\mathcal{F}^{\text{AdS}}} + e^{-\mathcal{F}^{\text{BH}}} + \dots$$

\longleftrightarrow
[Witten '98]

Does susy BH contribute to the gravitational path integral w/ susy boundary conditions?

$$\begin{aligned}\mathcal{I}_N(\tau) &= \text{Tr}_{\mathcal{H}_{\text{SYM}}} (-1)^F e^{2\pi i \tau n} e^{-\beta H^{\text{susy}}} \\ &= \int_{S^3 \times_{\tau} S^1} DA_{\mu} D\lambda \exp(-S_{\text{SYM}}(A_{\mu}, \lambda)) \\ &= \int_{\partial(\text{AdS}) = S^3 \times_{\tau} S^1} Dg_{\mu\nu} D\psi_{\mu} \exp(-S_{\text{grav}}(g_{\mu\nu}, \psi_{\mu}))\end{aligned}$$

- Euclidean BH: infinite throat (in the interior of AdS)
Temperature and chemical potentials all frozen!
⇒ Gibbons-Hawking procedure is singular

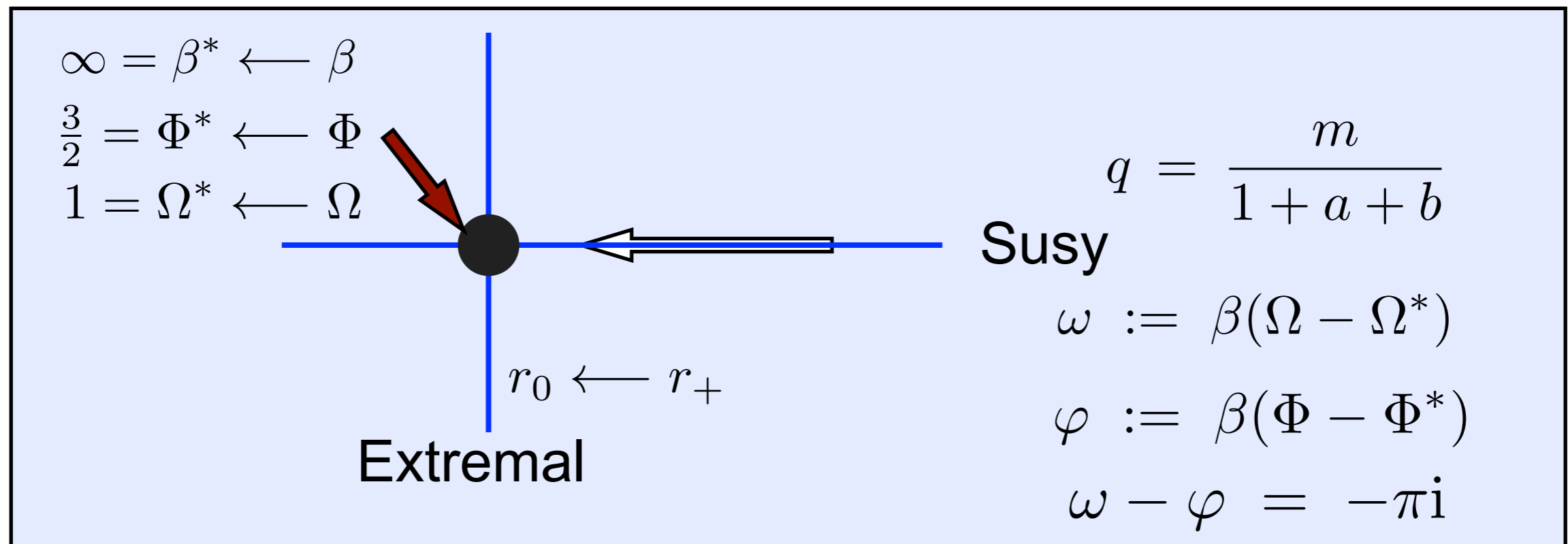
Deform susy extremal BHs to susy non-extremal solutions, ...

- Minimal 5d sugra

$$\mathcal{L} = R + \Lambda - F^2 + A \wedge F \wedge F$$

- Four-parameter family of BH solutions: $(a = b, m, q)$

[Gutowski, Reall, '04] [Chong, Cvetič, Lu, Pope, '05; Kunduri, Lucietti, Reall, '06]



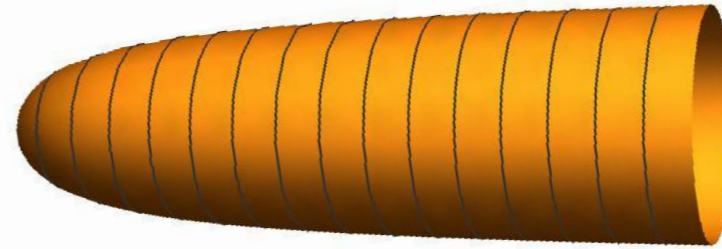
[cf Silva '06.]

[Cabo-Bizet, Cassani, Martelli, S.M. '18]

... then take a limit towards the susy BH

- Smooth Euclidean geometry (complex!)

[See Aharony, Benini, Mamroud, Milan, '21]



$$\oint A^R = \pi i$$

Note: rules for complex solutions not clear, but this clarifies a bit why index is compared to BH entropy.

- Killing spinor (with R-charge = 1) \Rightarrow susy preserved in bulk and on boundary
- Gibbons-Hawking procedure now well-defined. BH contribution to index = (limit of) on-shell action.

- Action is β -independent

$$\omega = 2\pi i \tau$$

$$\mathcal{F}^{\text{BH}} = \frac{\pi i N^2 (2\tau + 1)^3}{27 \tau^2}$$

[cf Hosseini, Hristov, Zaffaroni '17 where this function was noted to be the Legendre transform of the BH entropy]

On-shell action of AdS BH = Free energy of matrix model? Other phases?

$$\mathcal{I}_N(\tau) = \sum_{n \in \mathbb{Z}} d_N(n) e^{2\pi i \tau n/3}$$

$$\log |d_N(n)| \stackrel{?}{\sim} N^2 s(n/N^2)$$

$$n \rightarrow \infty$$

fixed N

$$n, N \rightarrow \infty$$

fixed $j = \frac{n}{N^2}$



$$\log \mathcal{I}_N(\tau) \stackrel{?}{\sim} \frac{\pi i N^2 (2\tau + 1)^3}{27 \tau^2}$$

$$\tau \rightarrow 0$$

fixed N

$$N \rightarrow \infty$$

fixed τ

Other
singularities?

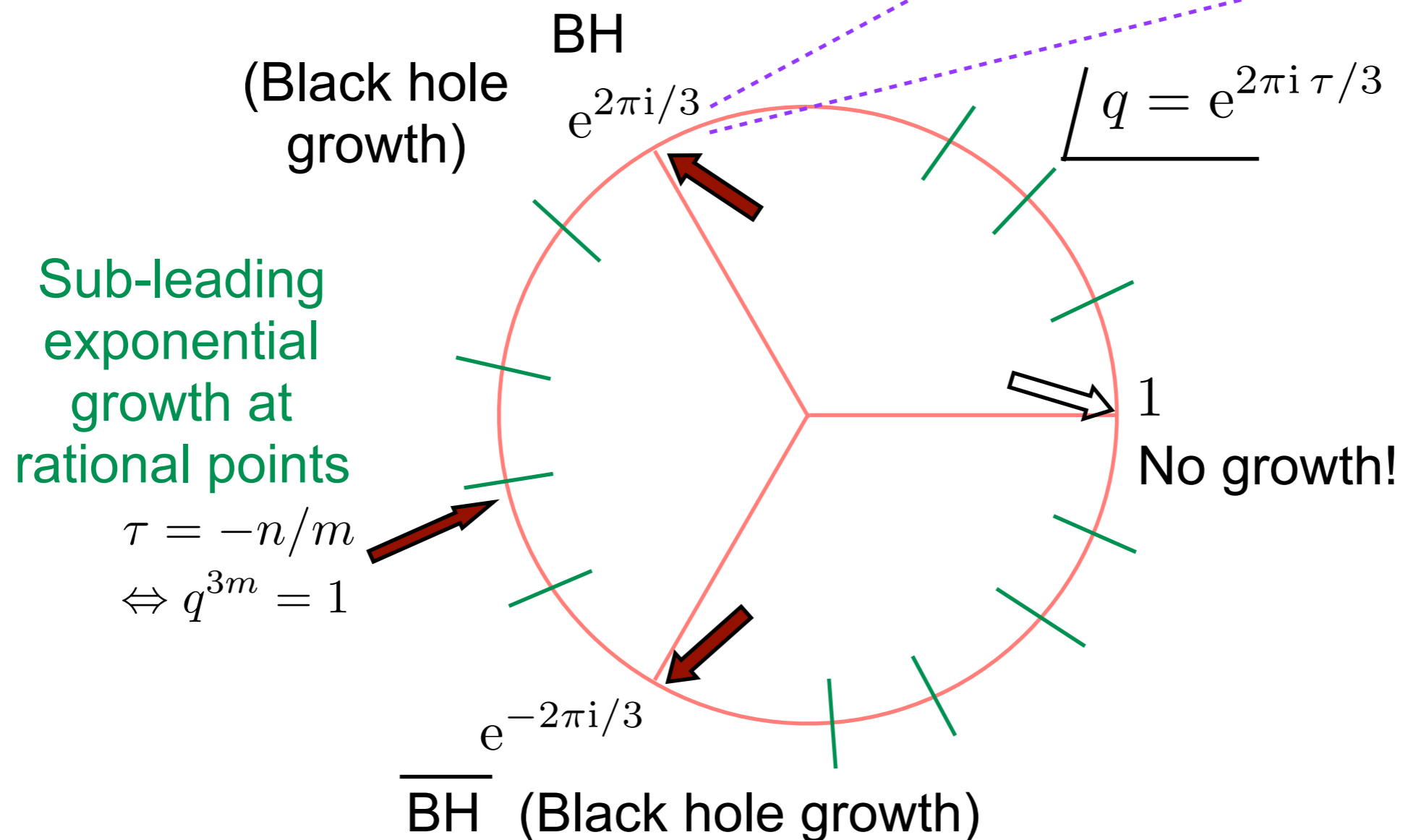
Cardy-like limit

Large-N limit

Saddles lie close to singularities of the index labelled by rational points/roots of unity!

$$\mathcal{I}_N(\tau) \simeq \sum_{\alpha \in \{\text{Saddles}\}} e^{-S_{\text{eff}}(\alpha; \tau)}$$

$$\tau \rightarrow \tau + 1, \\ \text{then } \tau \rightarrow 0$$



(Consistent w/
bulk-bndry
dictionary)

[Cabo-Bizet,
Cassani, Mar-
telli, S.M. '18]

1. Perturbation theory around each saddle is calculated from asymptotic analysis

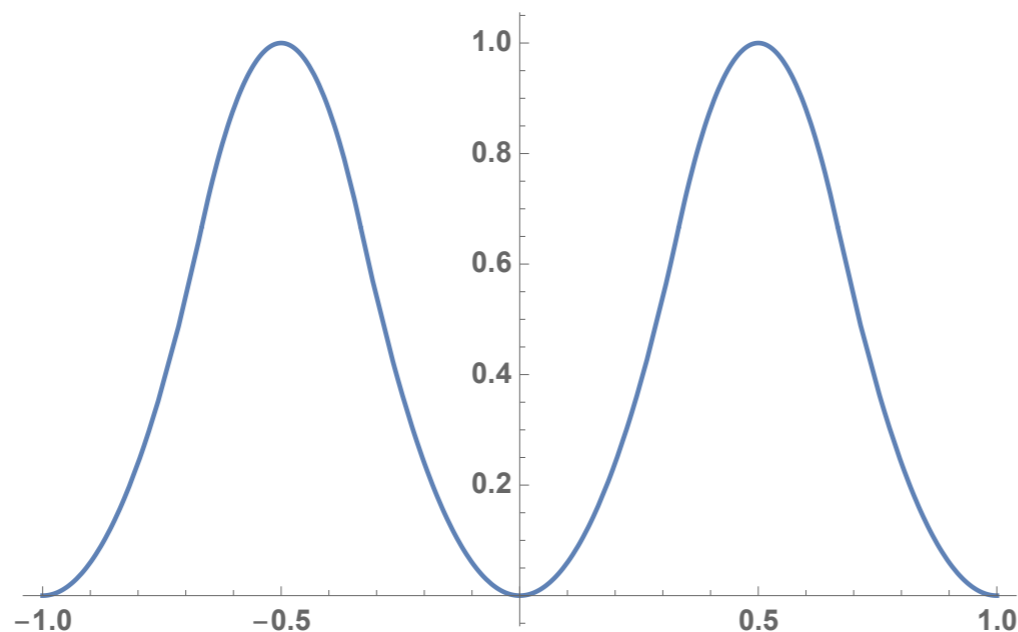
$$\mathcal{I}_N(\tau) = \int [D\underline{u}] \exp\left(-\sum_{i,j=1}^N V_\tau(u_i - u_j)\right)$$

$$e^{-V_\tau(z)} = \gamma_e(z + 2\tau; \tau) \times \gamma_e\left(z + \frac{2}{3}\tau; \tau\right)^3$$

As $\tau \rightarrow 0$,

$$-\frac{1}{2\pi i} \text{Log} \gamma_e(z + r\tau) \sim \frac{1}{6\tau^2} \bar{B}_3(z) + \frac{1}{6\tau} (r-1) \bar{B}_2(z) + \dots$$

Periodic Bernoulli polynomials



[Kim, Kim, Song '19]

[Cabo-Bizet, Cassani, Martelli, S.M. '19]

[Flavor: Honda '19; Ardehali, '19]

[Ardehali, Hong, Liu, '19], [Goldstein + Jejjala, Lei, van Leuven, Li '20, '21], ...

[Gonzalez-Lezcano, Hong, Liu, Pando-Zayas '21]

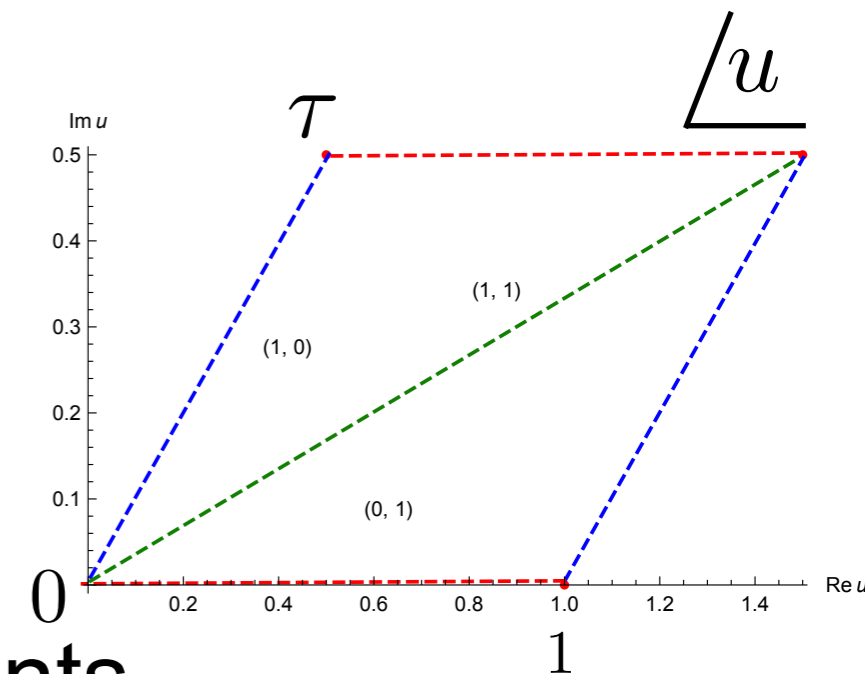
All-orders

[Ardehali, S.M. '21]

2. Relation to number-theoretic function allows for control at large-N

[S. Bloch, D. Wigner '78]

- Doubly-periodic extension of $V_\tau(u)$ to complex plane (torus) allows us to easily find saddle-point solutions.



Family of saddles labelled by lattice points

$$(m, n) \quad m\tau + n \in \mathbb{Z}\tau + \mathbb{Z} \quad \gcd(m, n) = 1$$

- Action of saddle calculated as the integral along a given direction is easy — use Fourier series (Bloch formula)!

[Cabo-Bizet, S.M. '19]

- Meromorphic (multi-valued) version and doubly-periodic (single-valued) version [S. Garoufalidis, S.M., D. Zagier, in progress]

[cf Method 3 (Bethe ansatz) Benini-Milan'18,...]. Answers agree when valid.

The perturbative action of the (m,n) saddle is a simple rational function of τ

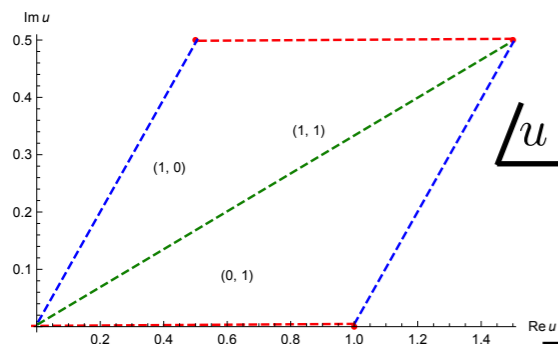
$$m \neq 0$$

$$S_{\text{eff}}(m, n; \tau) = \frac{N^2 \pi i}{27 m} \frac{(2\tilde{\tau} + \chi_1)^3}{\tilde{\tau}^2} + \text{const}, \quad \tilde{\tau} = m\tau + n$$

$$S_{\text{eff}}(0, 1; \tau) = 0$$

$m + n \pmod{3}$	0	1	2
χ_1	0	1	-1

Dirichlet character mod 3.



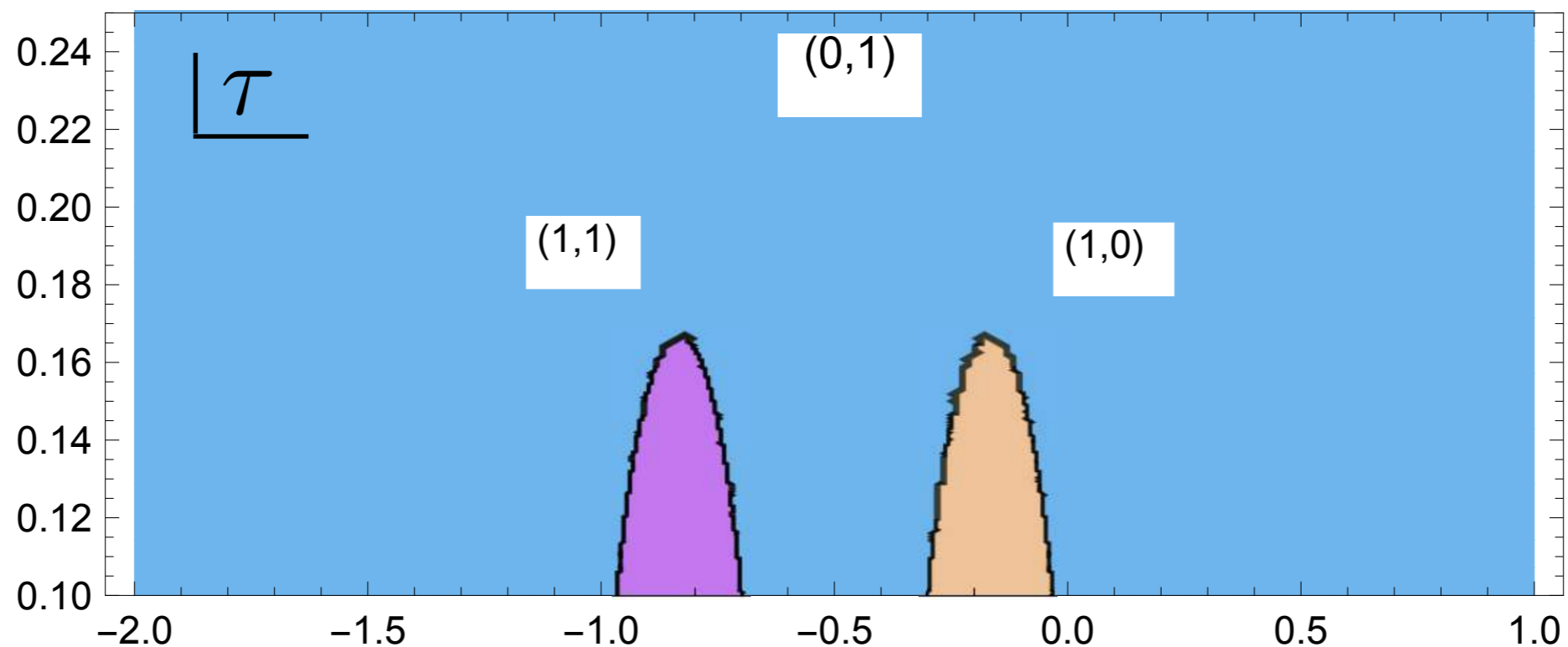
In other words, perturbation theory terminates!

In particular,

$$S_{\text{eff}}(1, 0; \tau) = \frac{N^2 \pi i}{27} \frac{(2\tau + 1)^3}{\tau^2} = \mathcal{F}^{\text{BH}}(\tau)$$

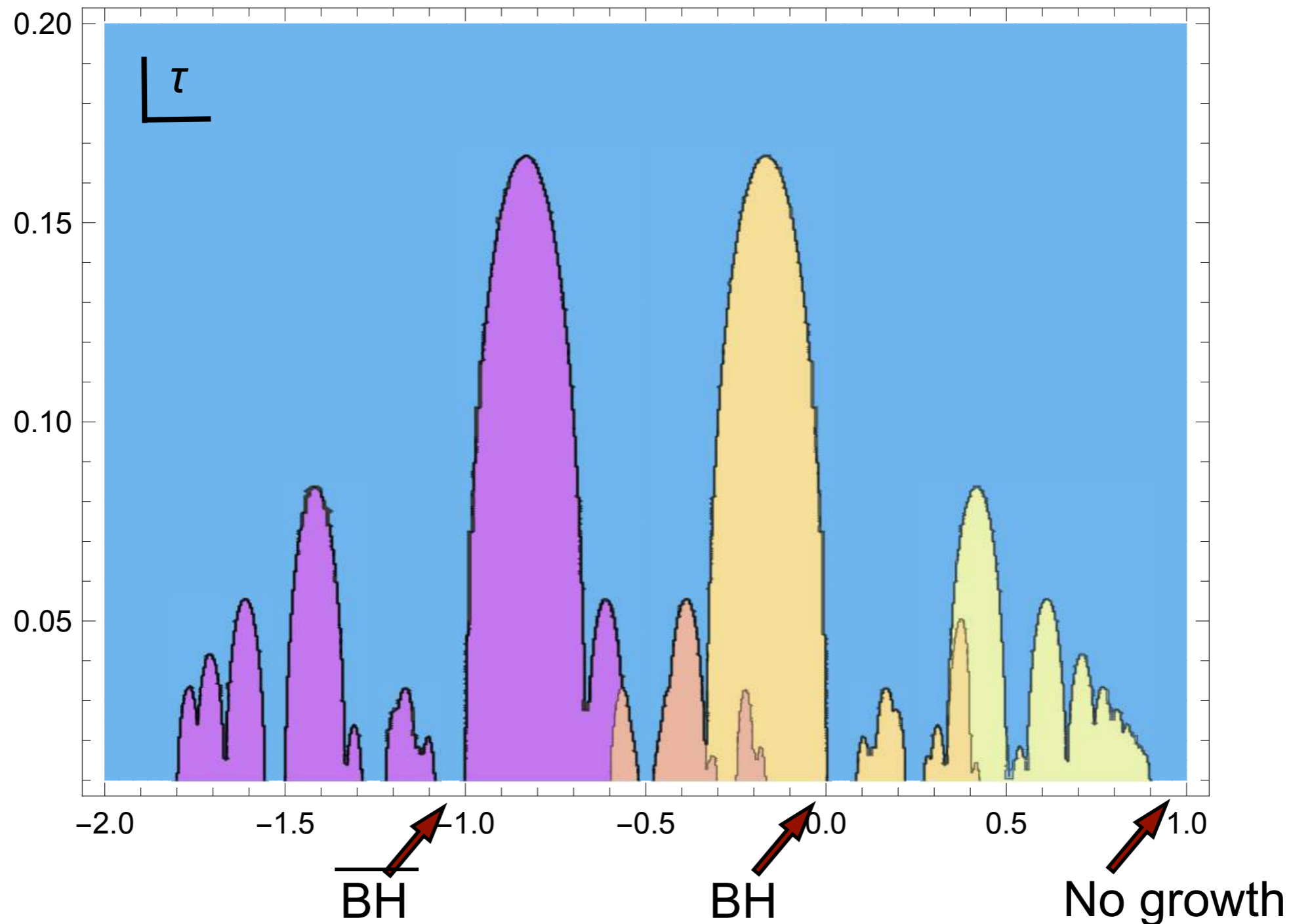
Phase structure: pure AdS dominates near $\tau = i\infty$, black hole dominates near $\tau = 0, \dots$

[Cabo-Bizet, S.M. '19]



... and new phases appear near rational points

[Cabo-Bizet, S.M. '19]



Comments on formula for index

$$\mathcal{I}_N(\tau) \simeq \sum_{m,n} \exp(-S_{\text{eff}}(m, n; \tau))$$

[Cabo-Bizet, S.M. '19]

[" + Cassani, Martelli '20]

[Cabo-Bizet '20]

- Complete asymptotics (perturbation exp) around each saddle. Only chamber-dependence remains as $N \rightarrow \infty$
- Prediction for family of gravitational solns (orbifolds)
See [Aharony, Benini, Mamroud, Milan, '21]
[cf AdS3 Maldacena-Strominger '98 and Dijkgraaf-Maldacena-Moore-Verlinde '00]
- Each term can be derived completely from 3d effective CS field theory! [Ardehali, S.M. '21 (m,n)] [Cassani, Komargodski '21 (1,0)]
[cf. Choi, Kim, Kim, Nahmgoong '18]

Things I said and would have liked to say

- BH states clearly present. Many phases. Full perturbative expansion understood.

- *Complex* saddles of matrix model and of gravity are important for full understanding of AdS/CFT.

- In each phase, perturbation theory explained by 3d effective Chern-Simons theory. [I could only touch upon this point.]

- Finite-N index interpolates between gravitons and BHs! cf. [Chang, Yin '13].
Index = sum over random partitions. cf. [Gross, Taylor '93]
[I did not have time to discuss this point.]

Thank you very much!